

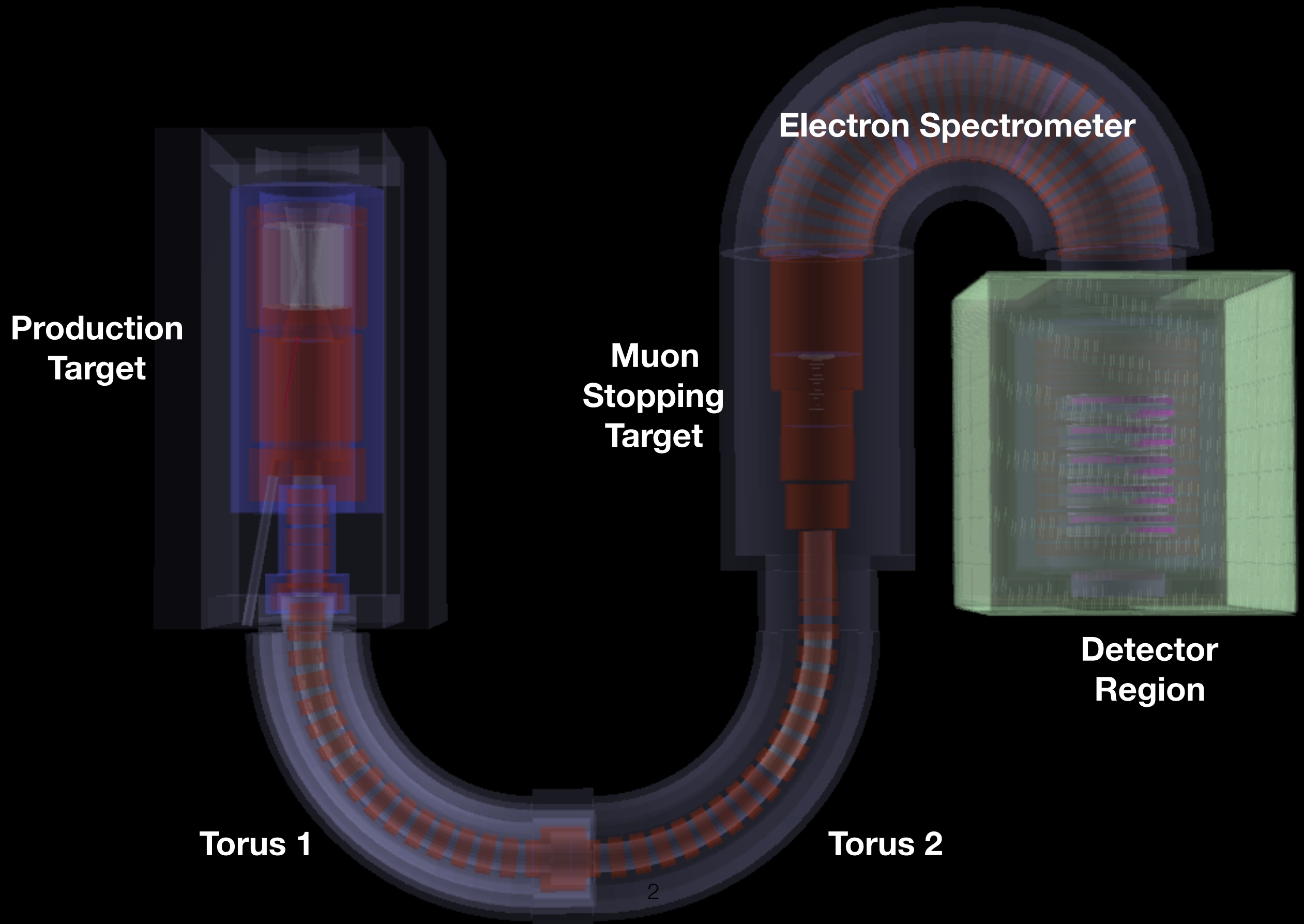


Stopping Target Optimization for COMET Phase II

Giorgi Kumsiashvili M2

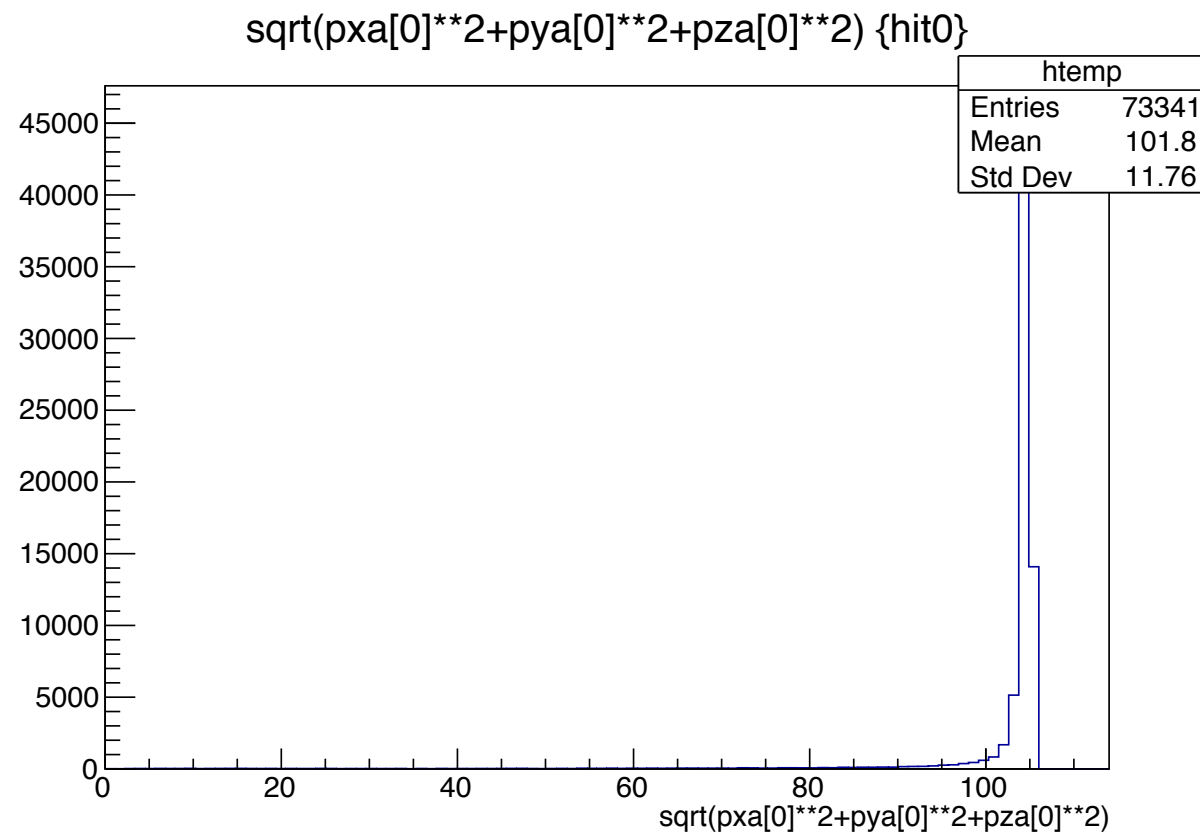
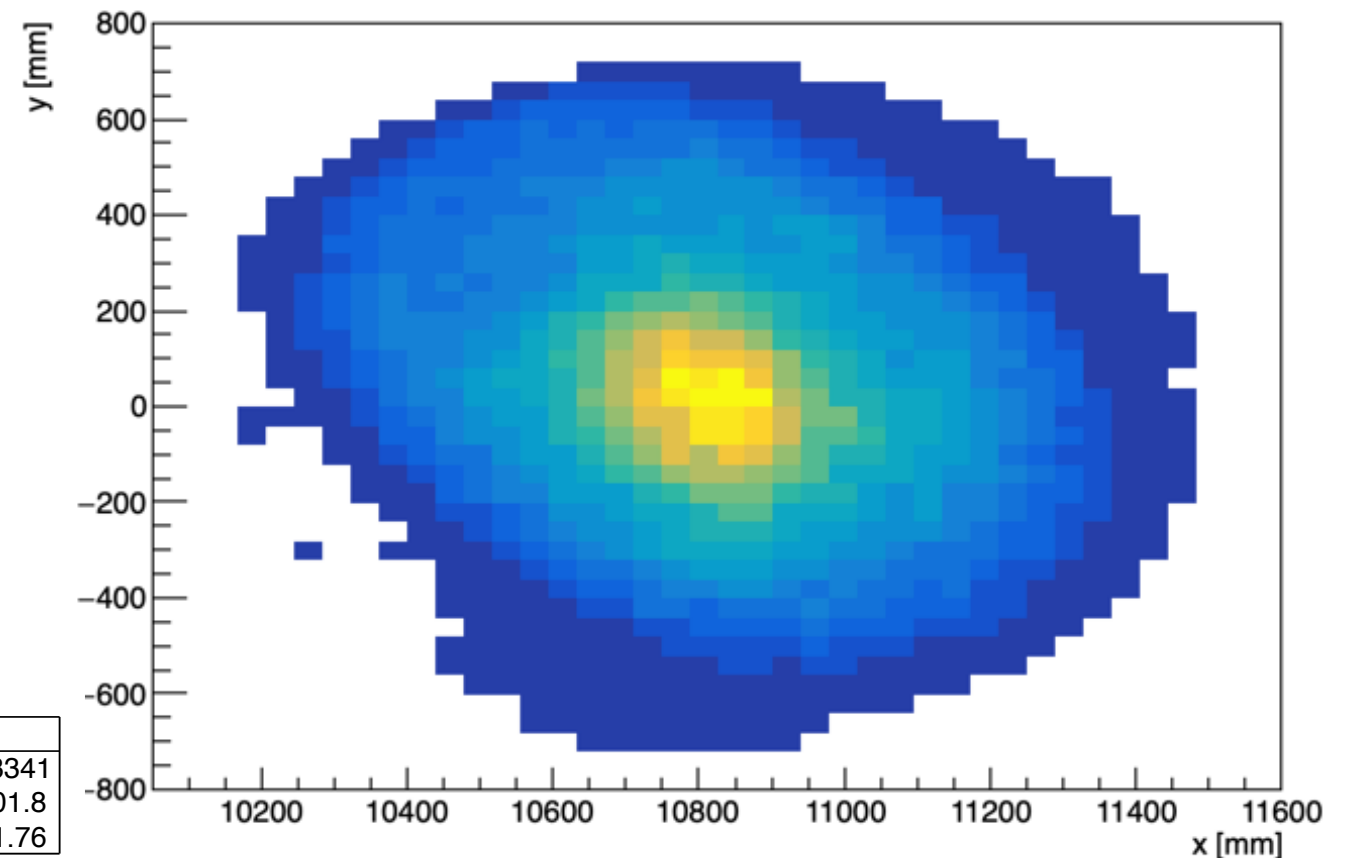
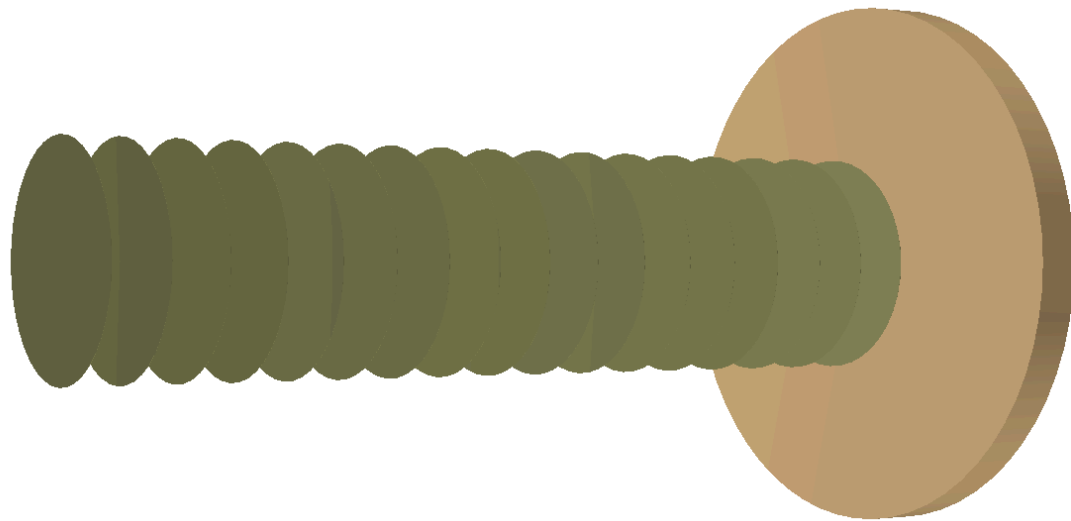
Year-end Presentation 2019

COMET Phase II Beamline



Original Target Design

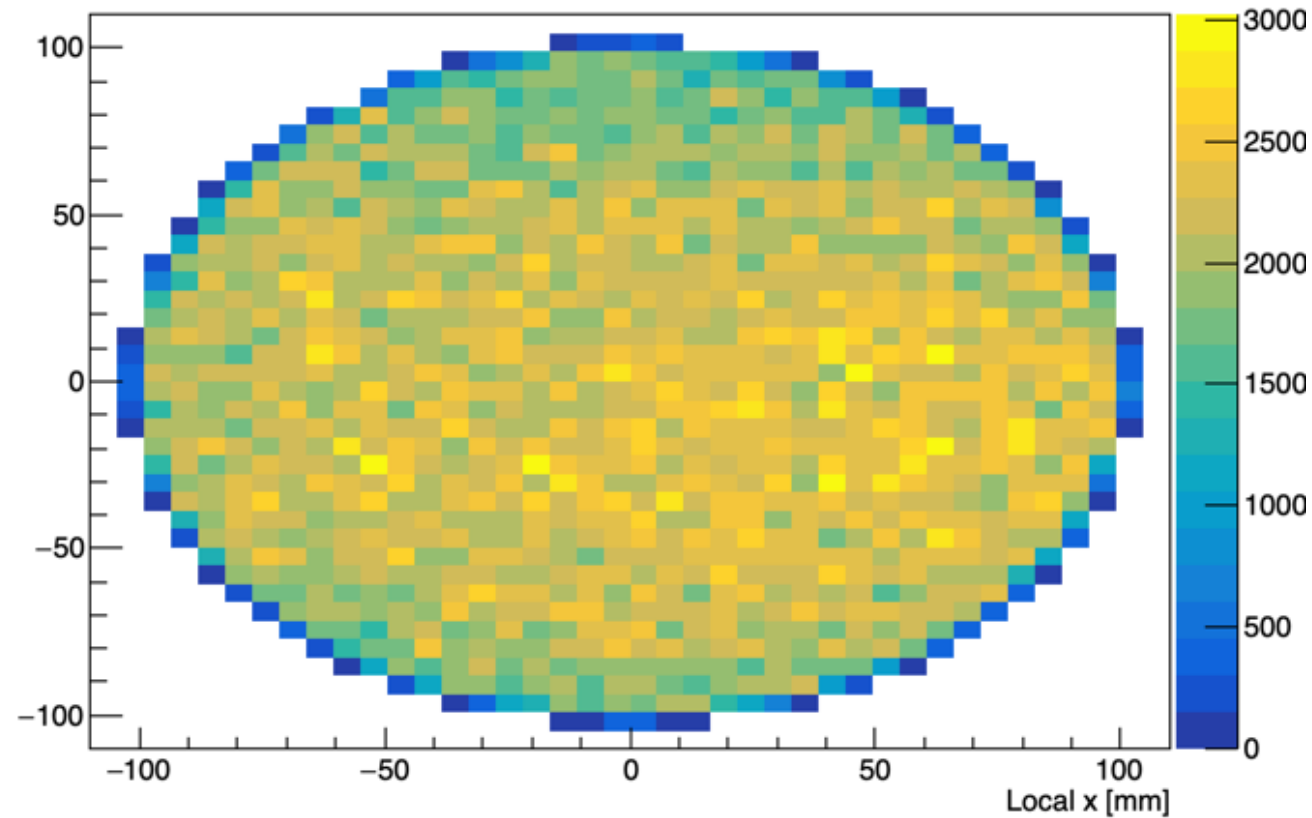
Signal Electron Distribution in Straw Tracker



Target Radius 100mm
Beam Blocker - Vacuum
DIO Blocker - Vacuum
Default Gradient Mag. Field

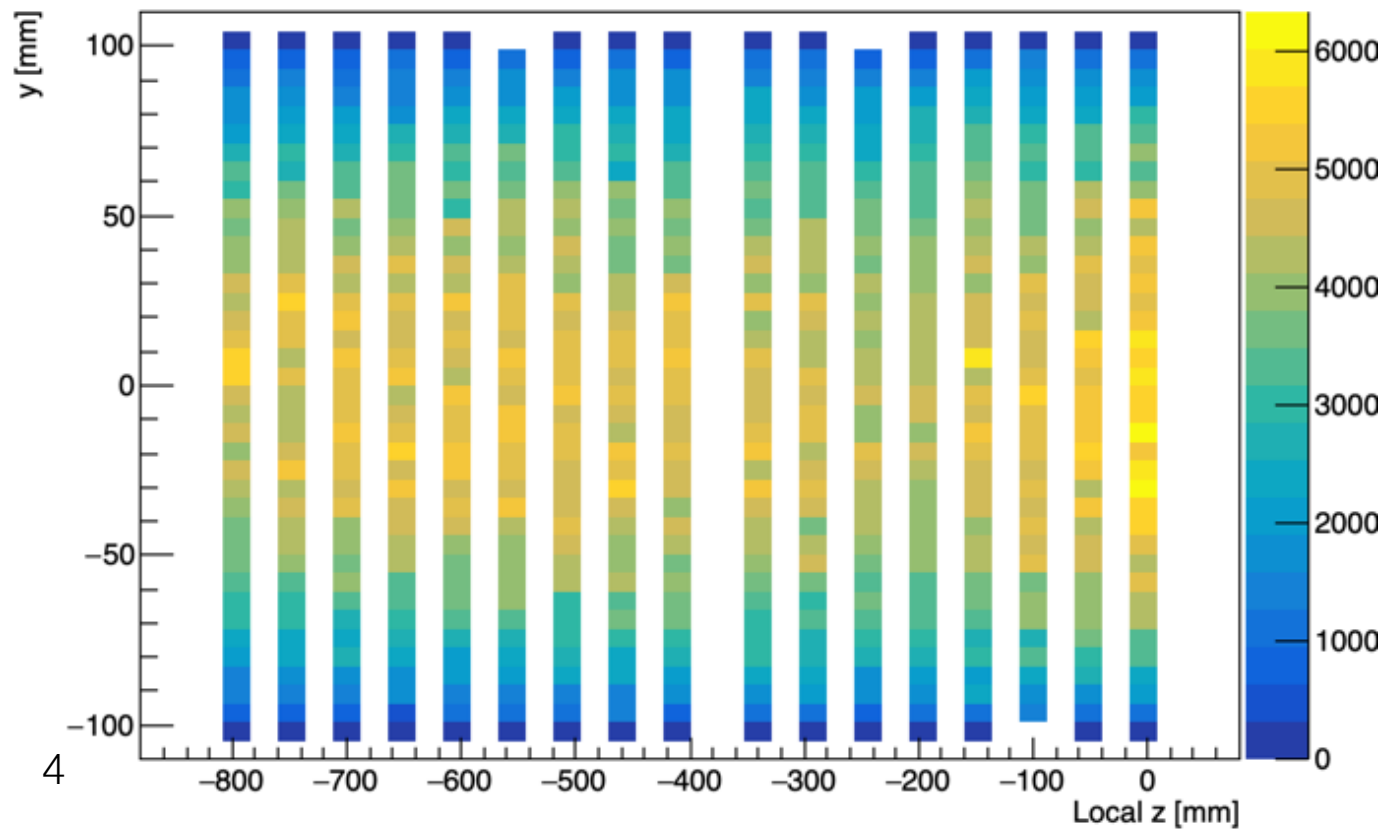
Signal Electrons

Initial Distribution of Signal Electrons



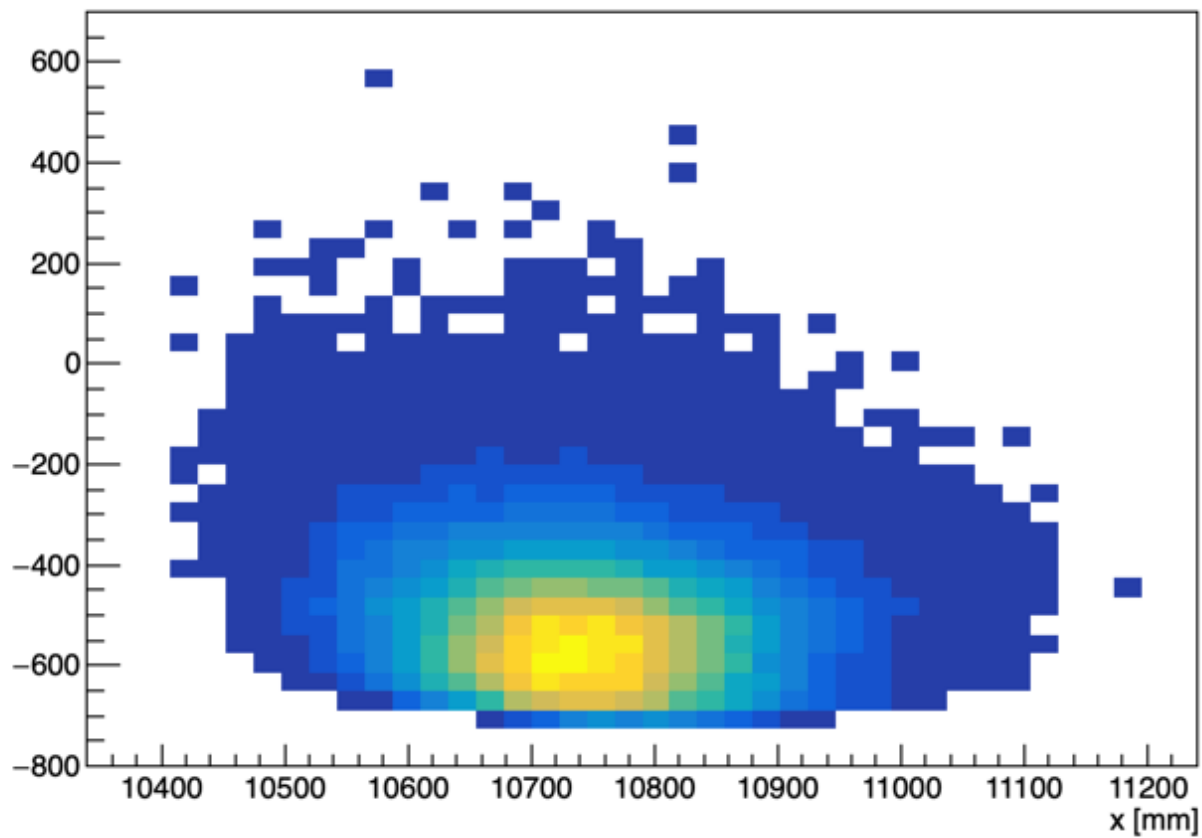
**Signal Electrons that are
later accepted by the detector**

Initial Distribution of Signal Electrons

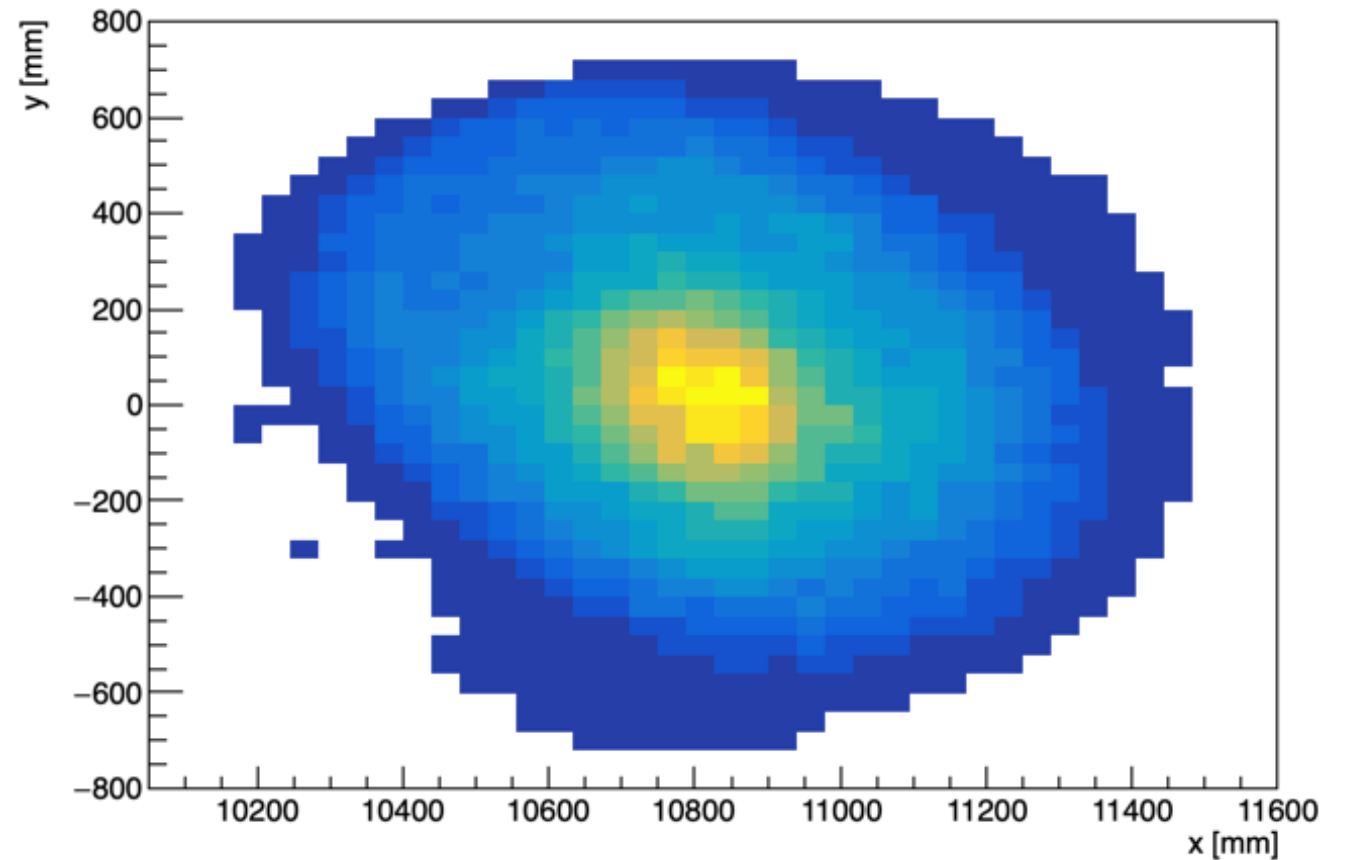


Distribution in Straw Tracker

DIO Hits in Straw Tracker



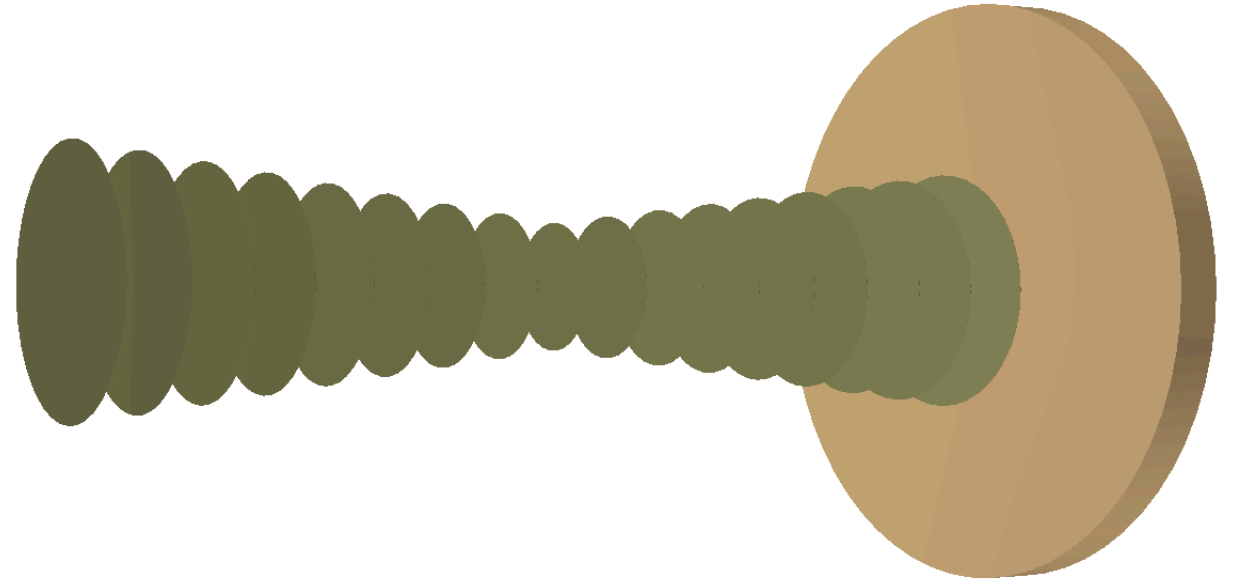
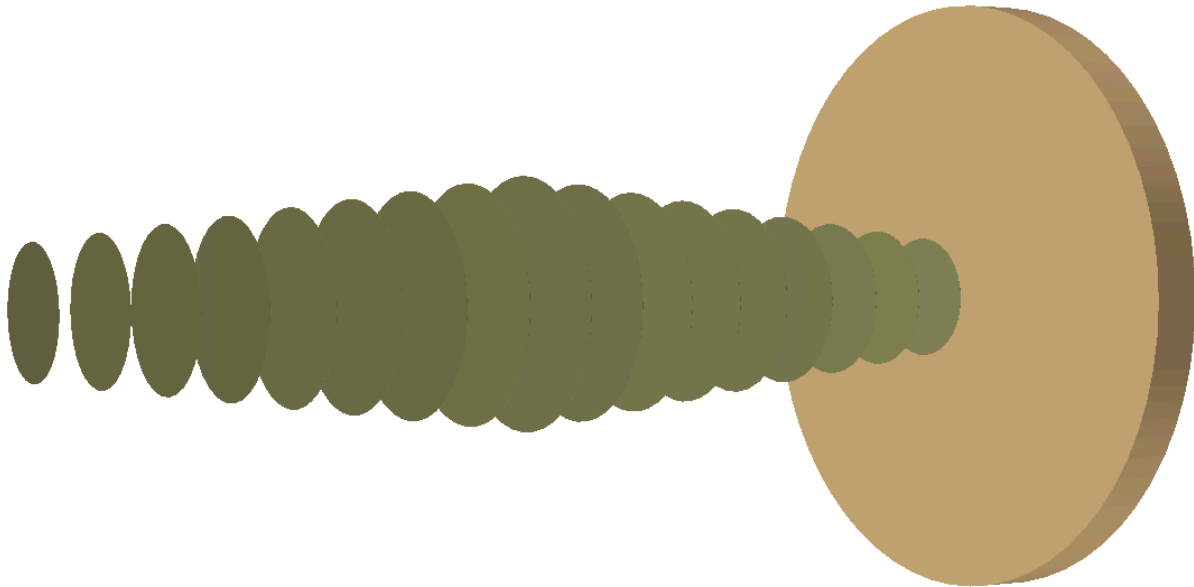
Signal Electron Distribution in Straw Tracker



Test Designs

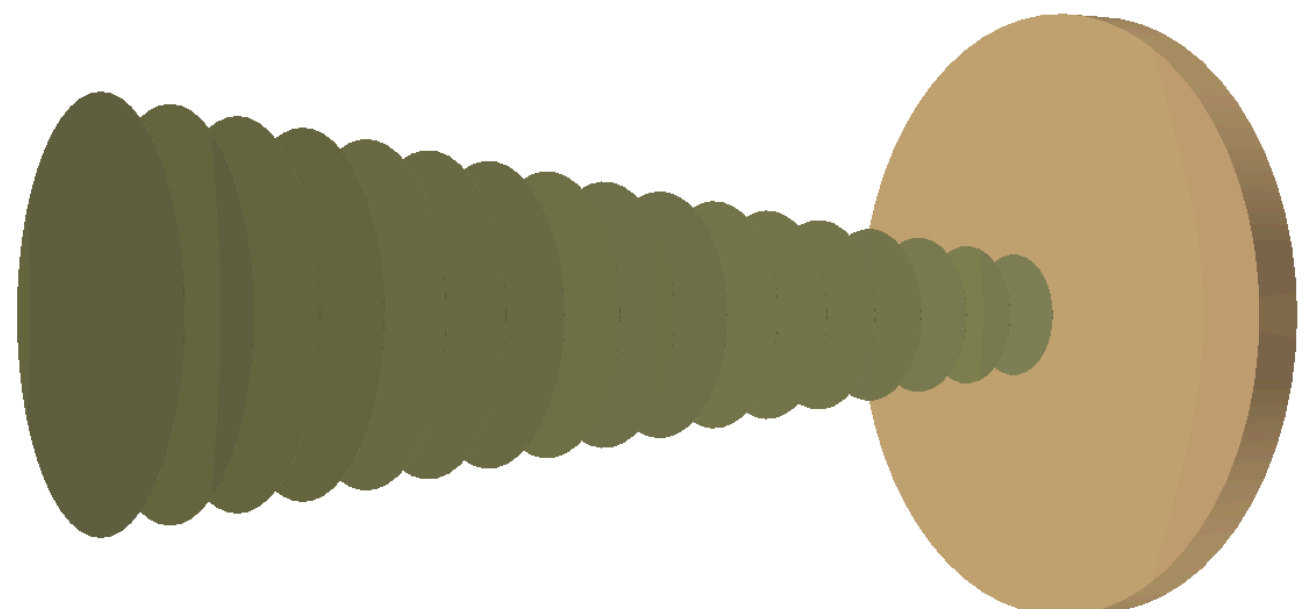
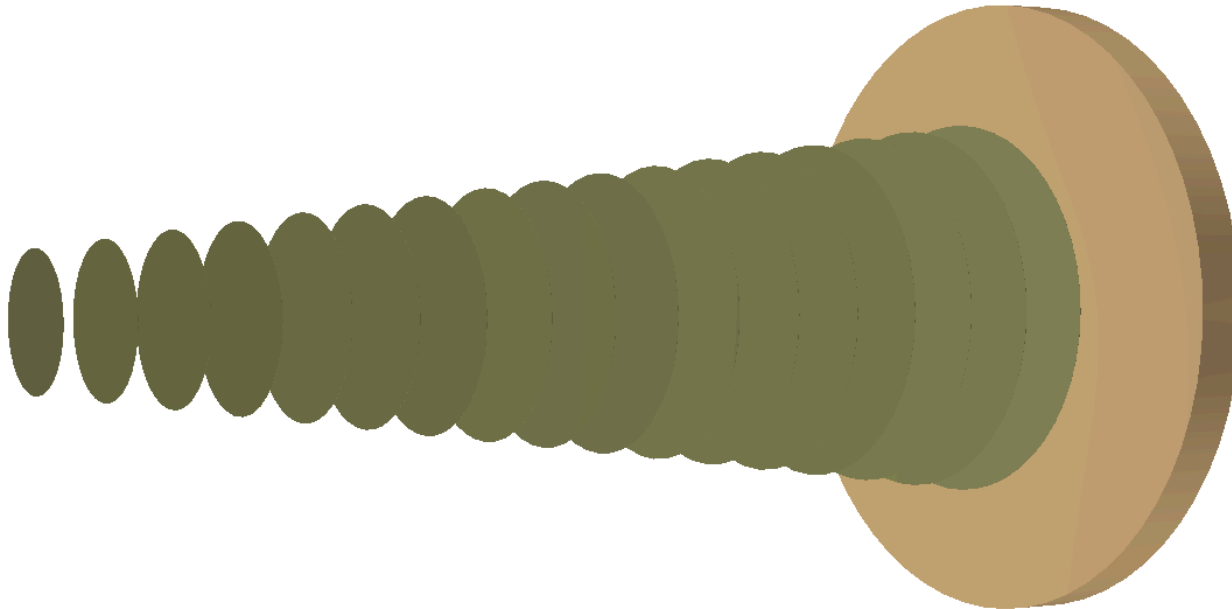
1st Disk - 50mm, 9th - 100mm, 17th - 50mm

1st Disk - 100mm, 9th - 50mm, 17th - 100mm



1st Disk - 50mm, 17th - 150mm

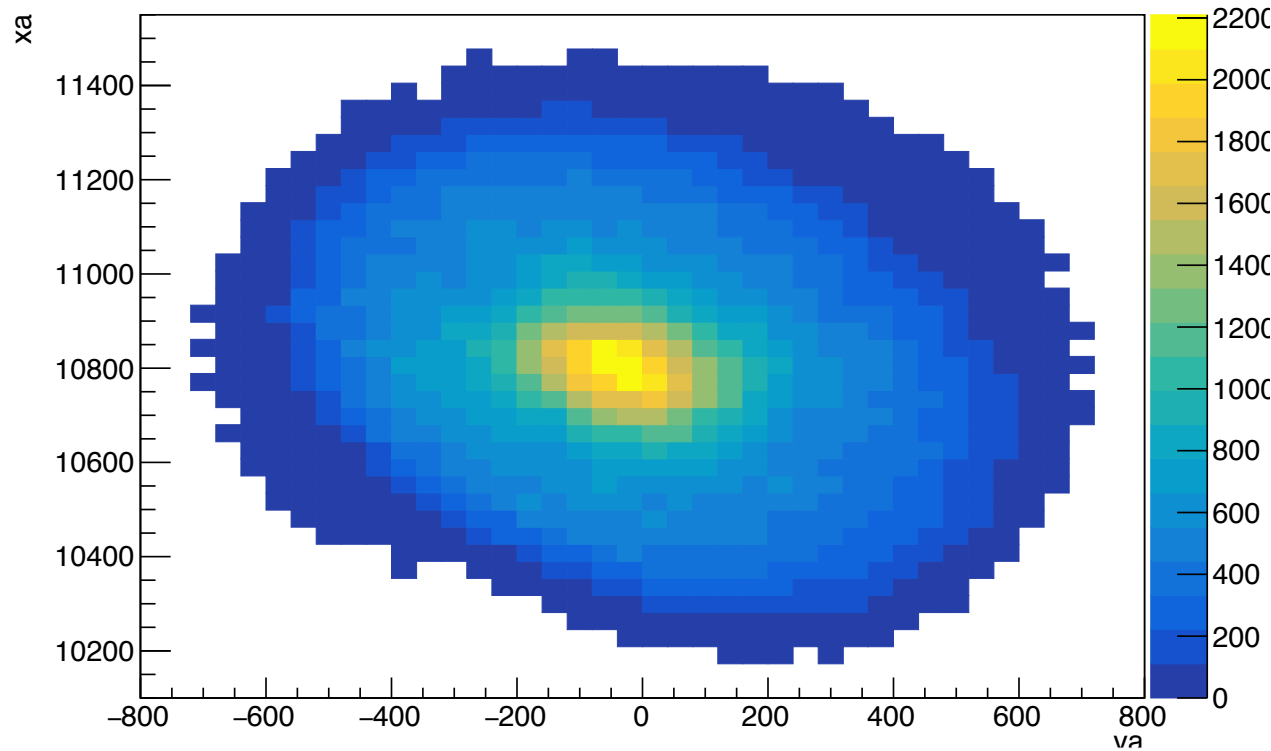
1st Disk - 150mm, 17th - 50mm



Distributions In Straw Tracker

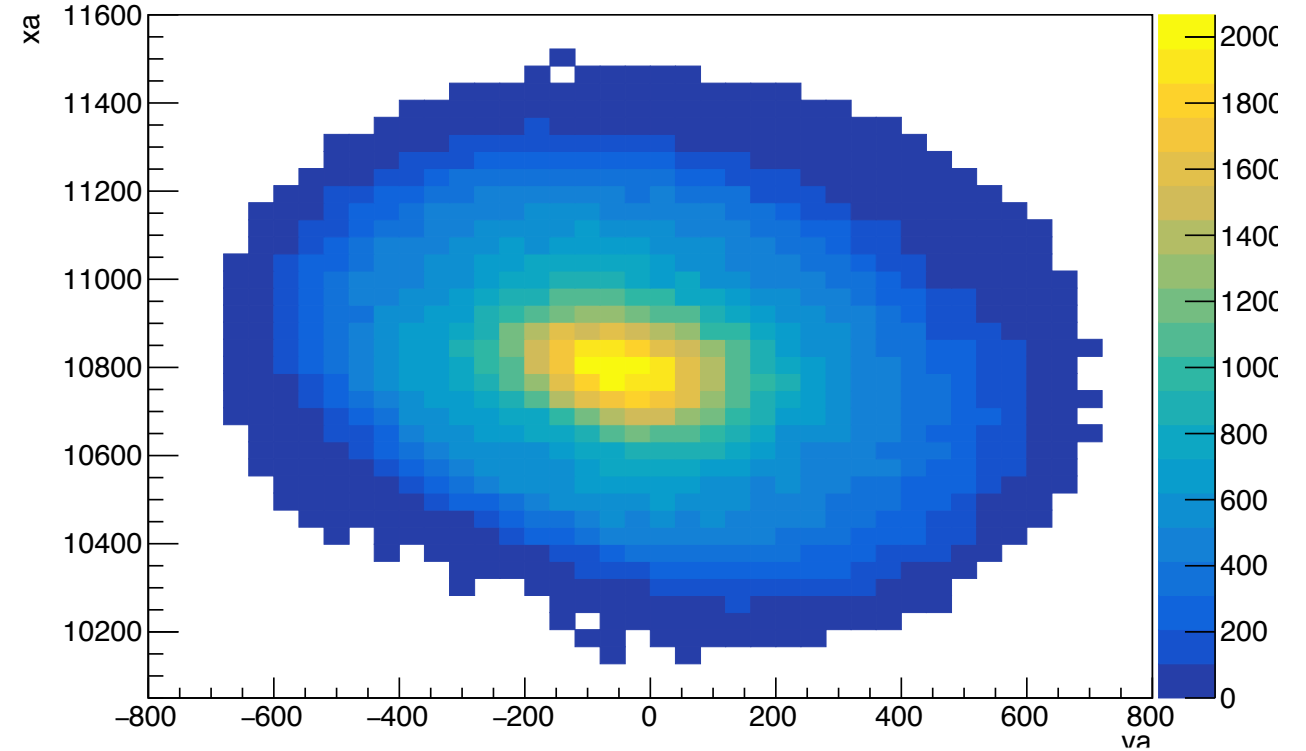
50mm-100mm-50mm

xa:ya {hit0&&hit1&&hit2&&hit3&&hit4&&xa>10000}



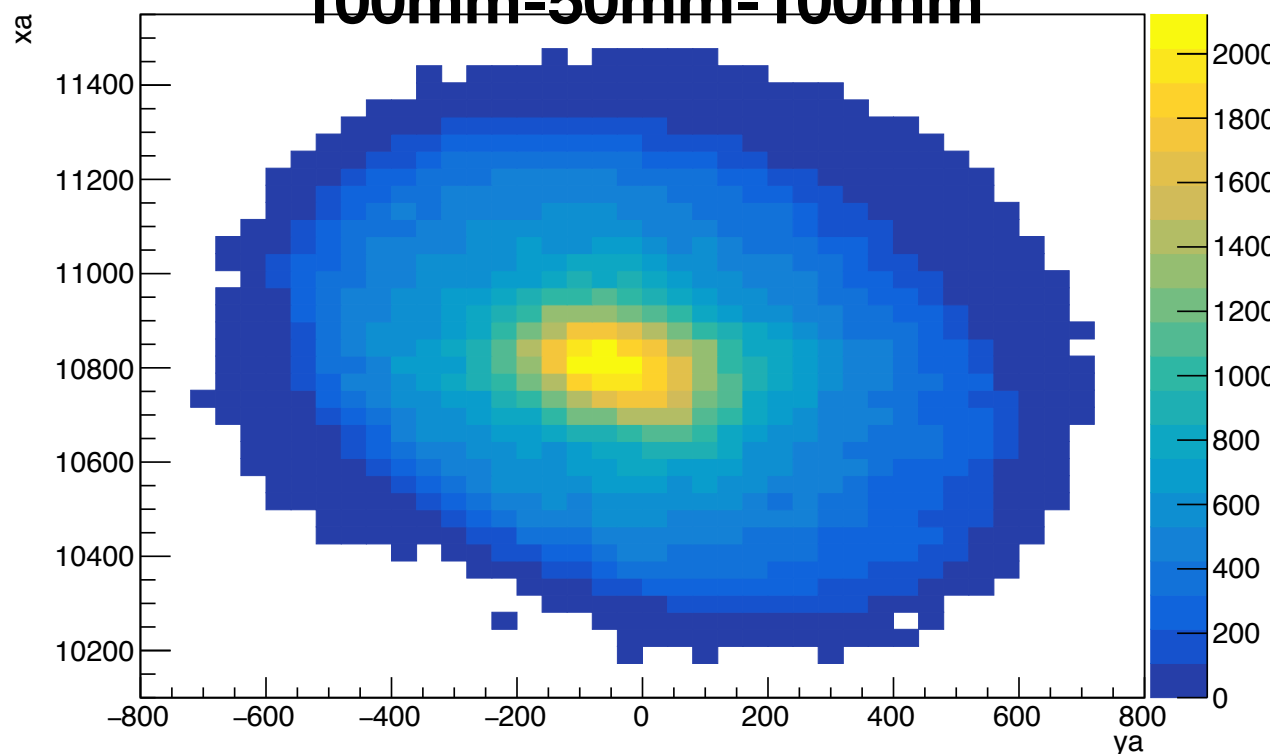
50mm-150mm

xa:ya {hit0&&hit1&&hit2&&hit3&&hit4&&xa>10000}



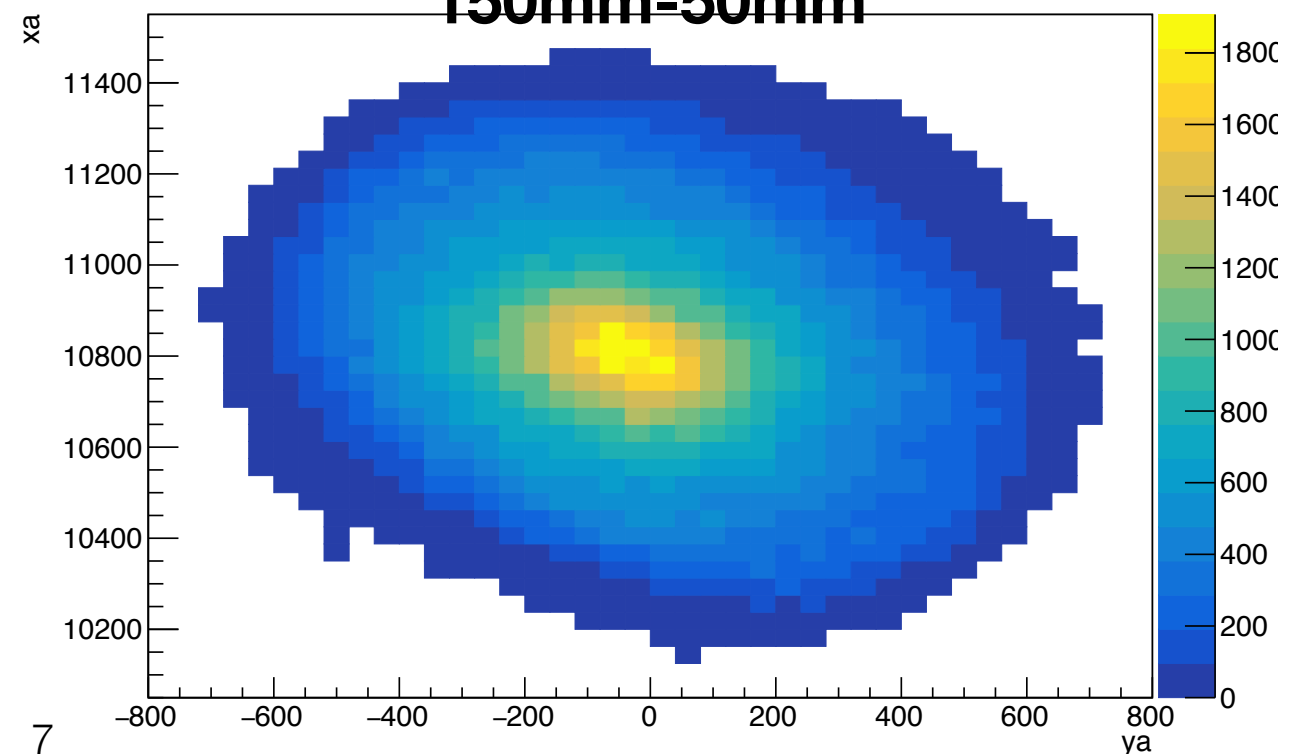
100mm-50mm-100mm

xa:ya {hit0&&hit1&&hit2&&hit3&&hit4&&xa>10000}



150mm-50mm

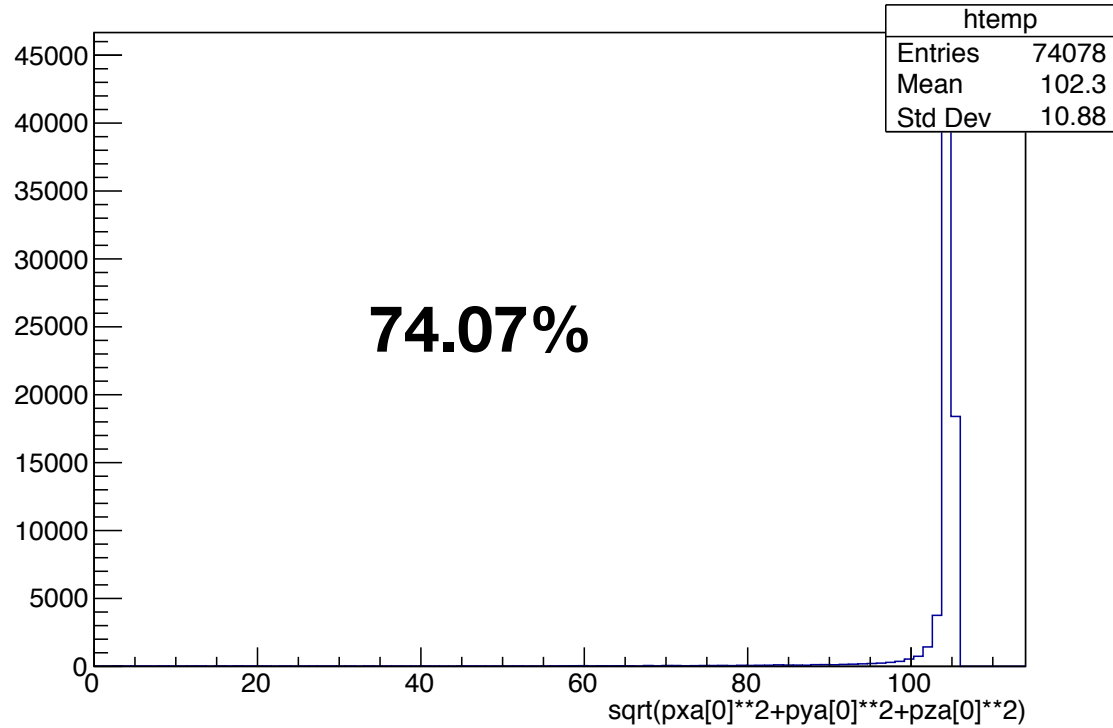
xa:ya {hit0&&hit1&&hit2&&hit3&&hit4&&xa>10000}



Momentum

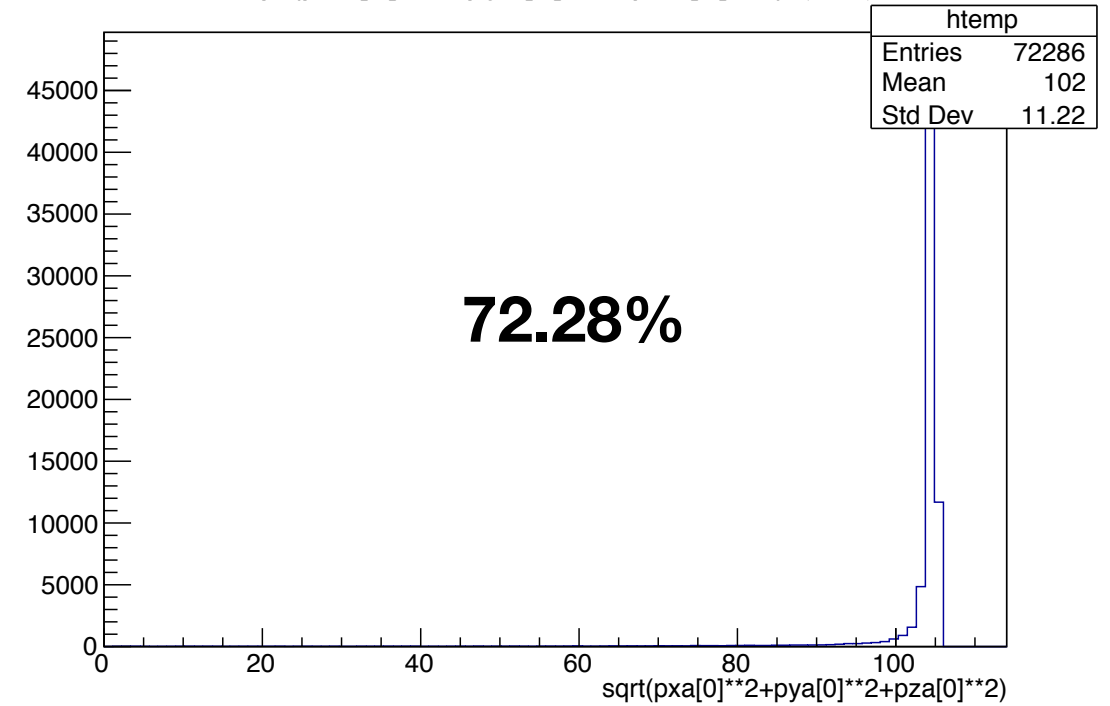
50mm-100mm-50mm

$\sqrt{pxa[0]**2 + pya[0]**2 + pza[0]**2}$ {hit0}



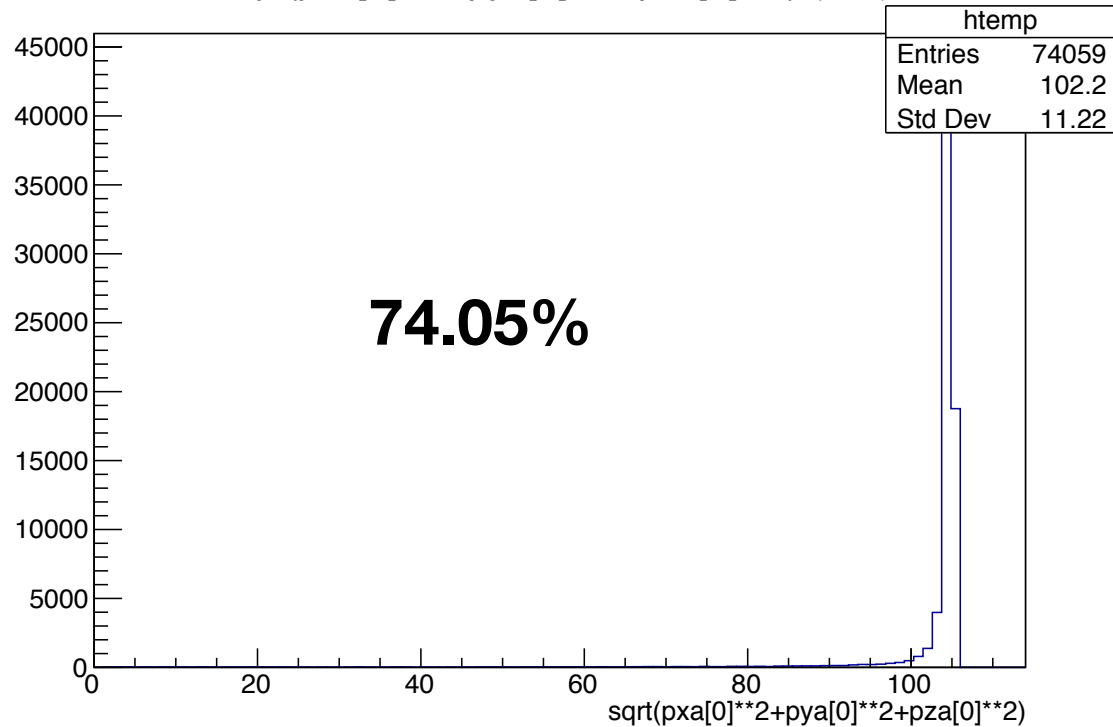
50mm-150mm

$\sqrt{pxa[0]**2 + pya[0]**2 + pza[0]**2}$ {hit0}



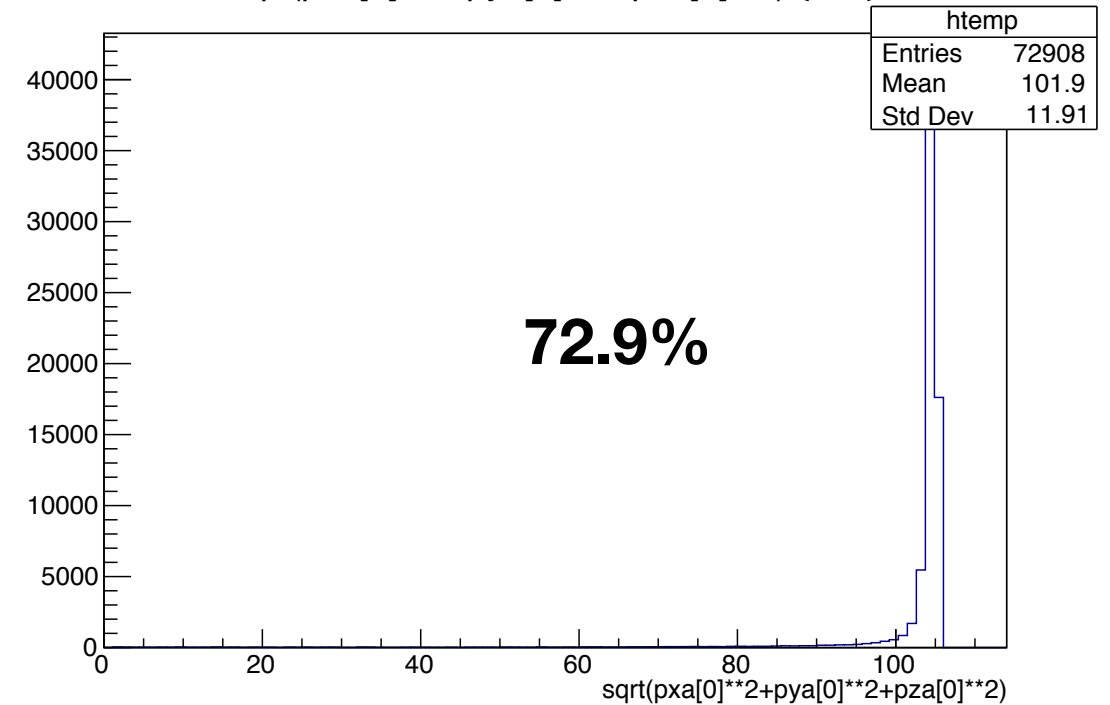
100mm-50mm-100mm

$\sqrt{pxa[0]**2 + pya[0]**2 + pza[0]**2}$ {hit0}

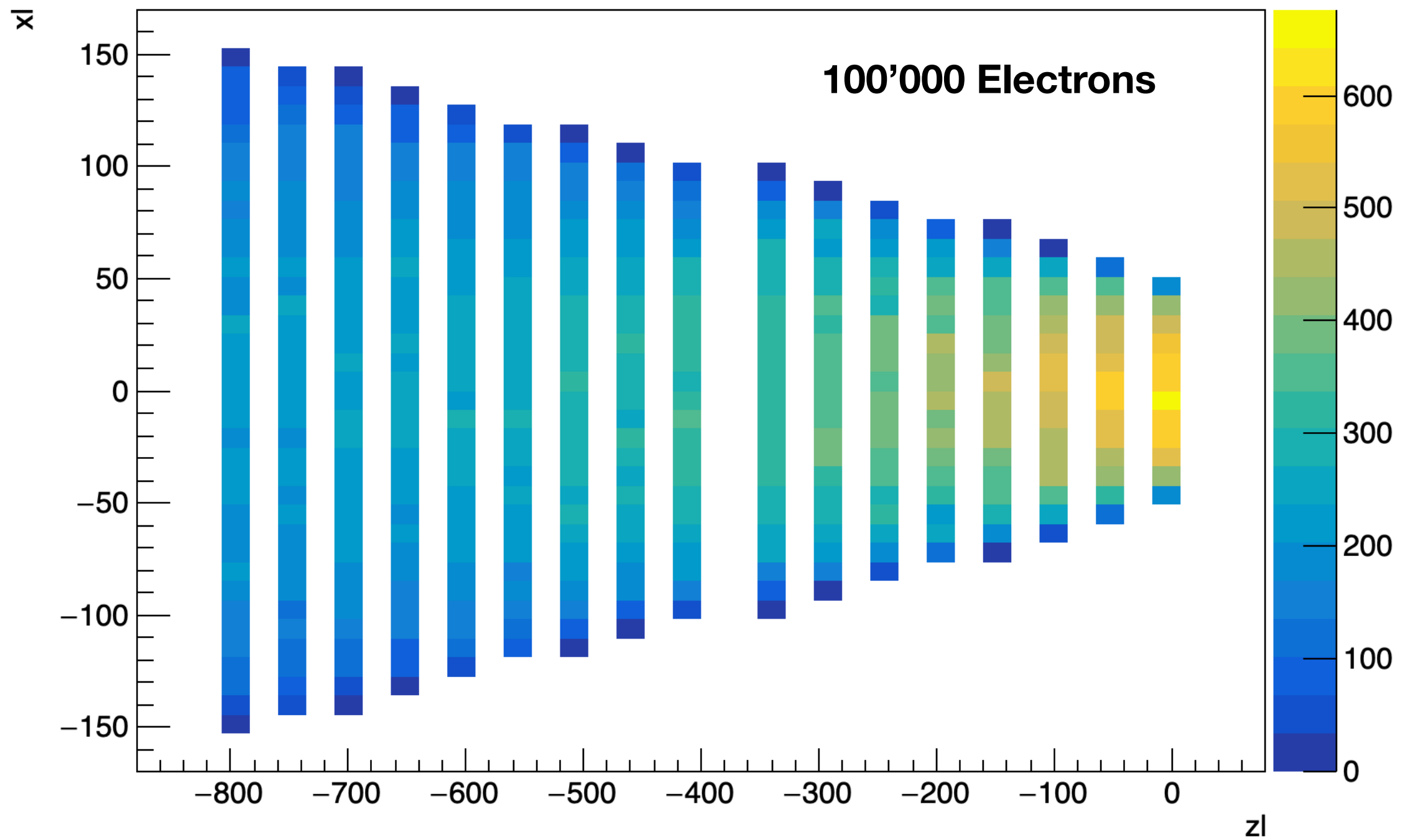


150mm-50mm

$\sqrt{pxa[0]**2 + pya[0]**2 + pza[0]**2}$ {hit0}

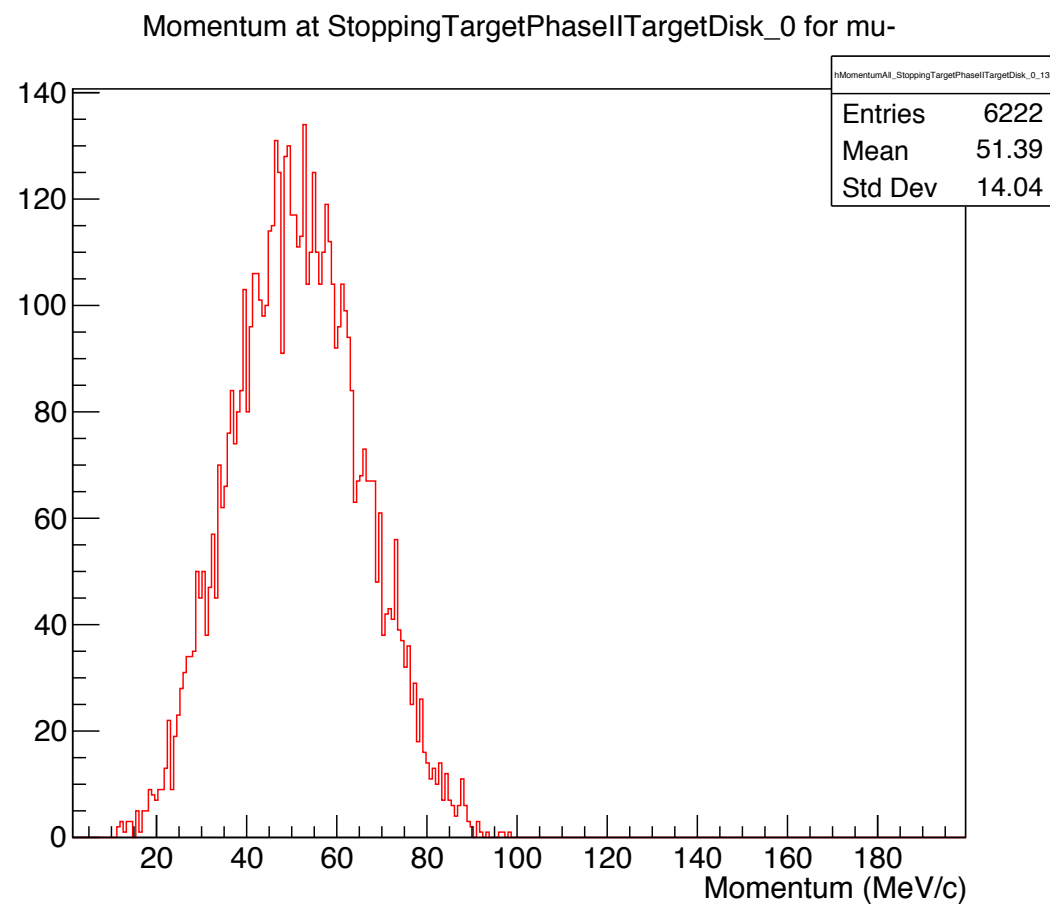


Initial Distribution

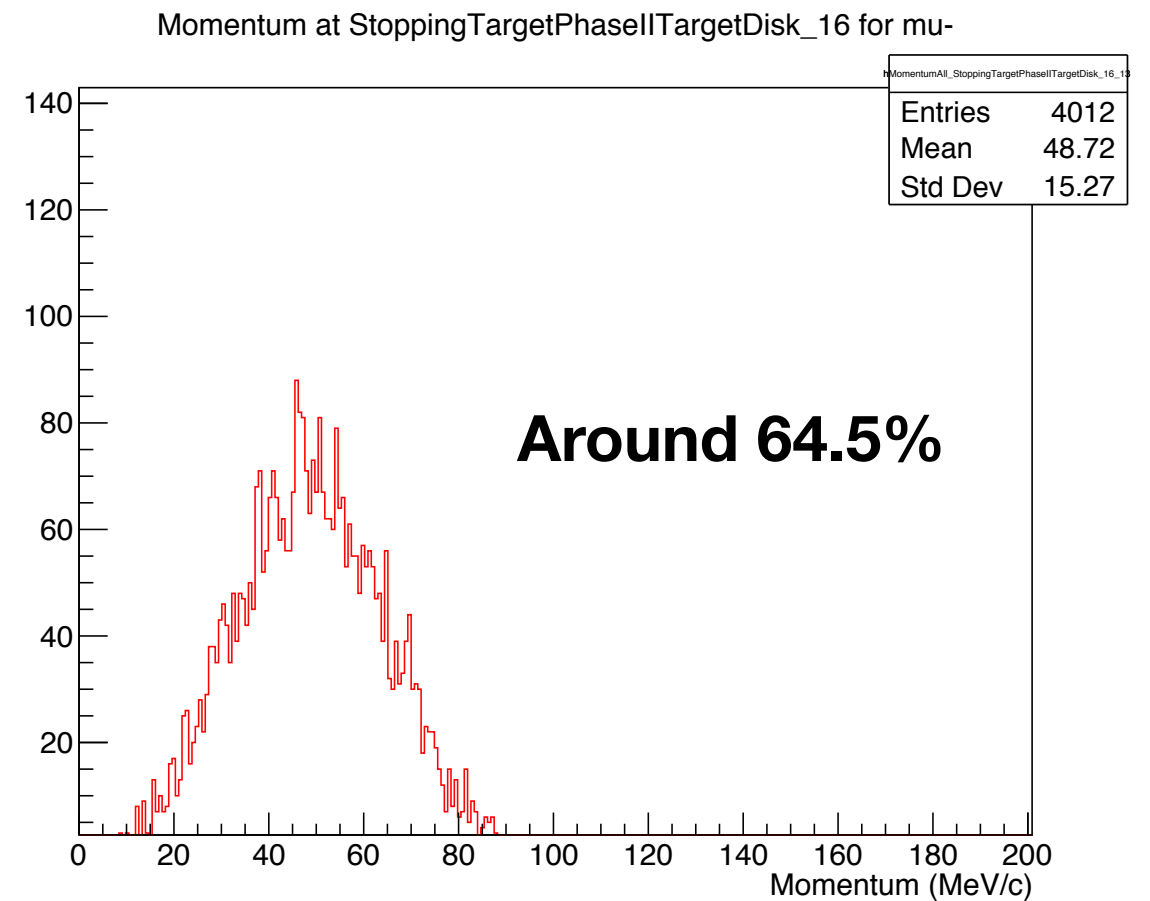


Mu- Momentum in Disks

In this example:
Number of disks - 17
Disk radius - 100mm



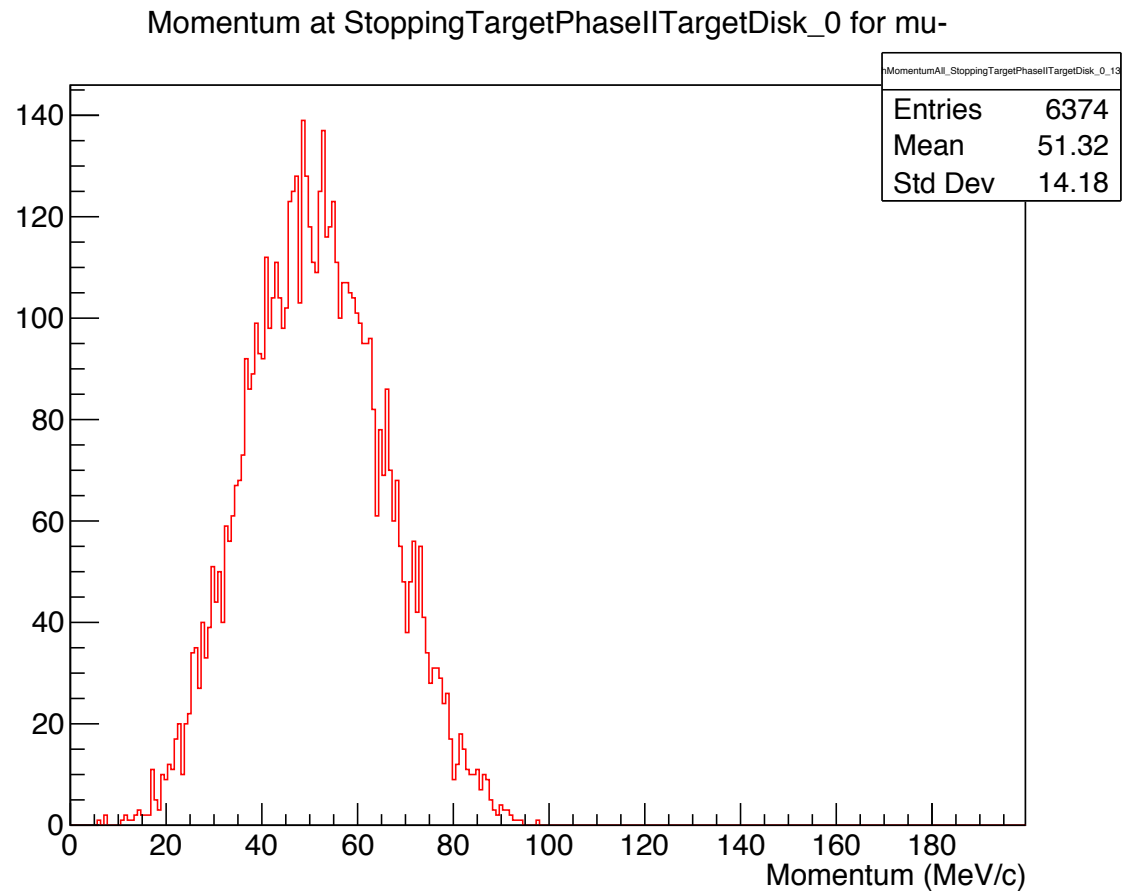
Disk 1



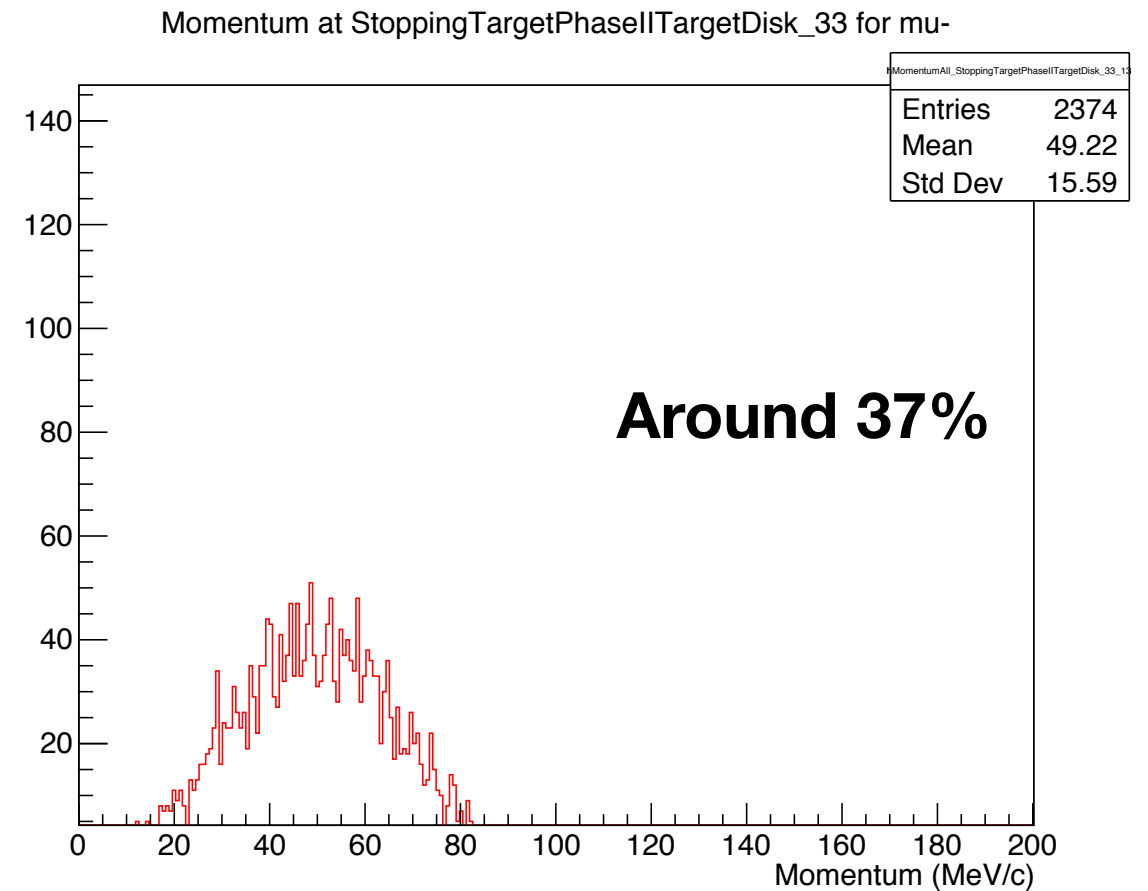
Disk 17

Mu- Momentum in Disks

In this example:
Number of disks - 34
Disk radius - 100mm



Disk 1



Disk 34

Around 37%

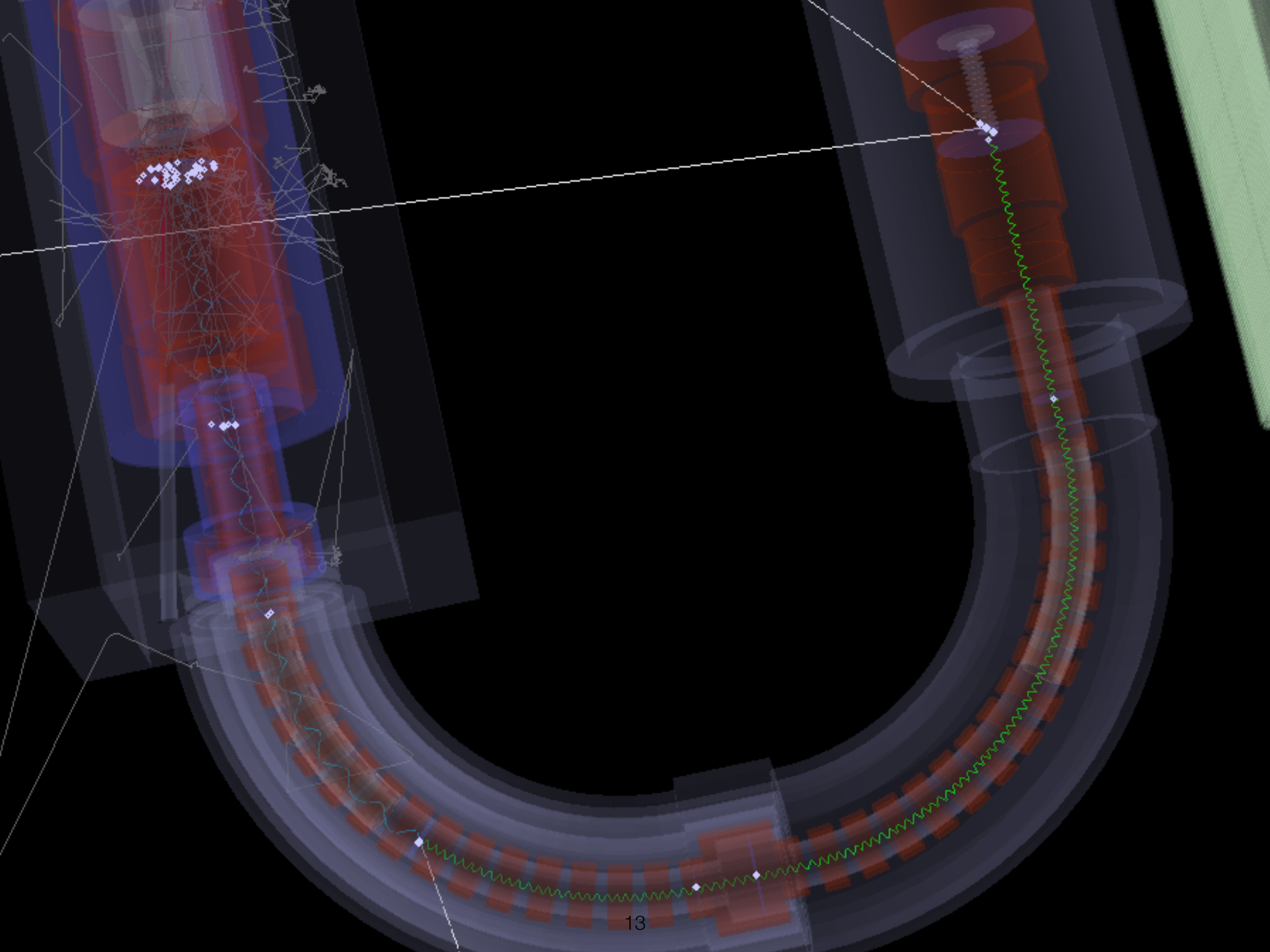
Summary and future work

- **The original muon stopping target design is not optimal.**
- **Test target designs yielded marginal improvements. There is a need to obtain realistic muon stopping distribution in target disks. (in progress)**
- **Significant number of Muons go through the target disks and downstream.**

What's next..

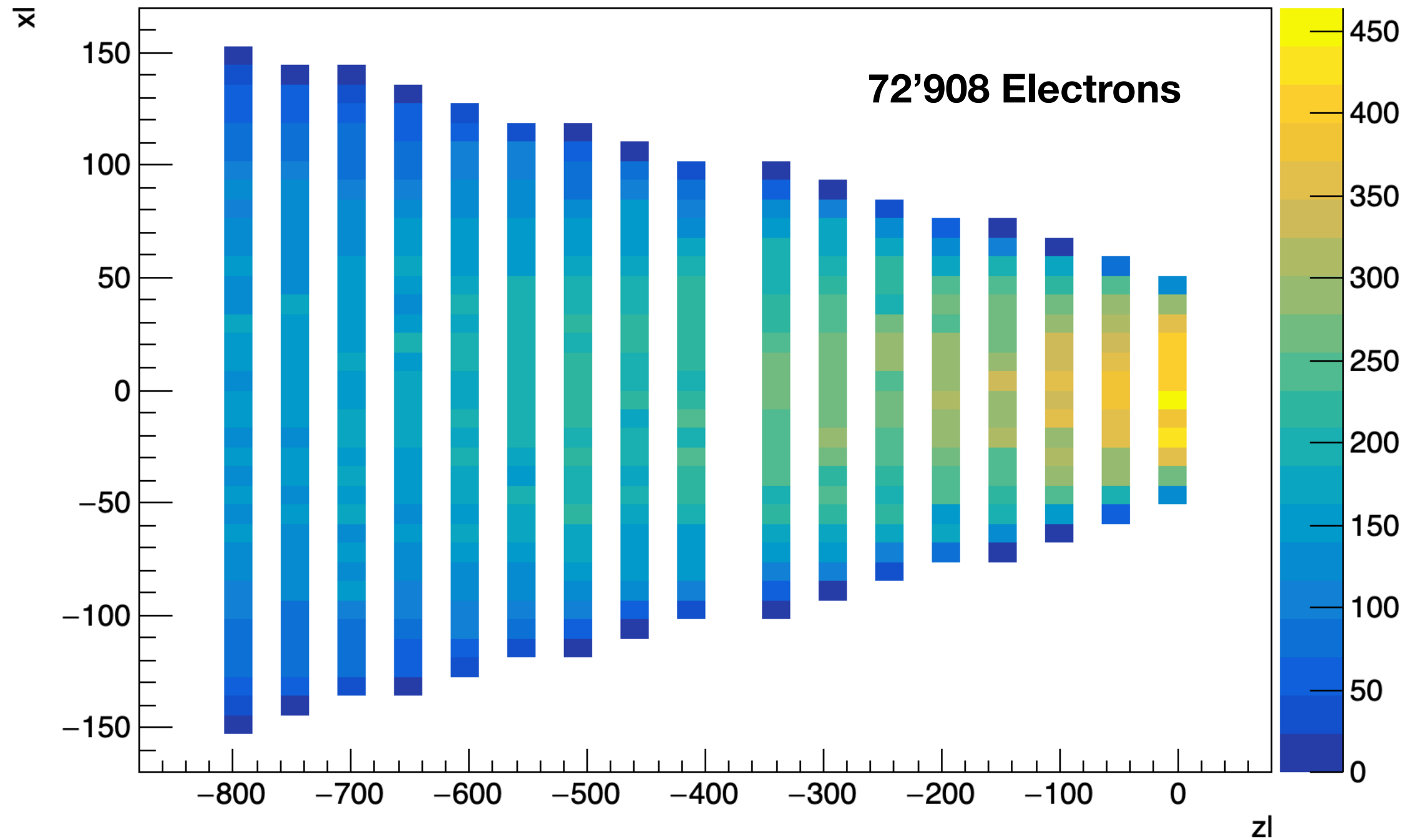
- **Run larger scale simulations for usage in signal acceptance tests.**
- **See if using 34 target disks hurt signal acceptance or not. (Smaller disks?.)**

Thank you for your attention!



BACKUP

Initial Distribution (That hit all 5 stations)



COMET Phase II

Background Estimations

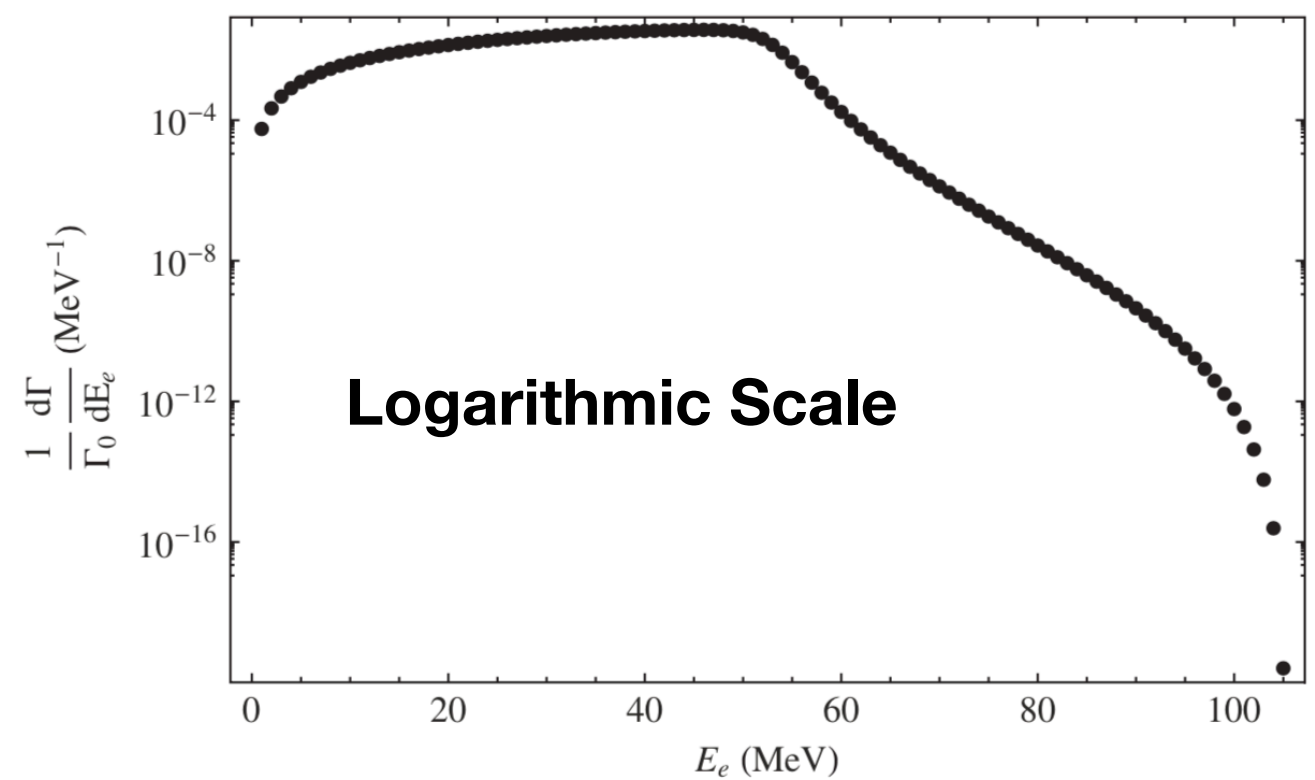
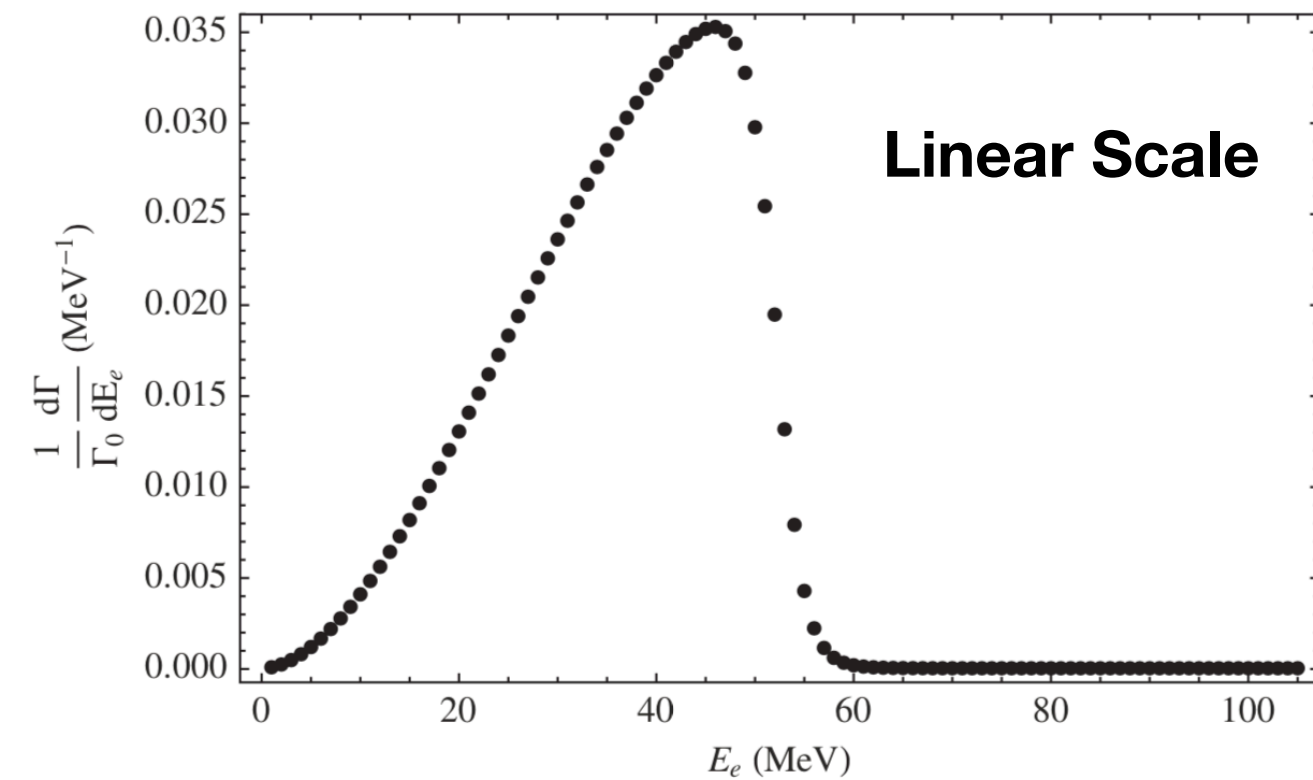
Type	Source	Ben's results	this study
Intrinsic	DIO	0.068	0.75
intrinsic	RMC	4.10×10^{-13}	2.46×10^{-12}
Delayed	RPC	1.18×10^{-6}	
Delayed	Beam	1.00×10^{-3}	
Delayed	Stopped $\bar{p}$	0.296	0.569
Delayed	π^- from $\bar{p}$	1.33×10^{-9}	
Prompt	RPC	1.23×10^{-3}	
Prompt	Beam	1.91×10^{-3}	
Prompt	π^- from $\bar{p}$	2.24×10^{-8}	
Cosmic rays		0.294	0.562
Total		0.662	1.9

From COMET Phase II note

Background estimations

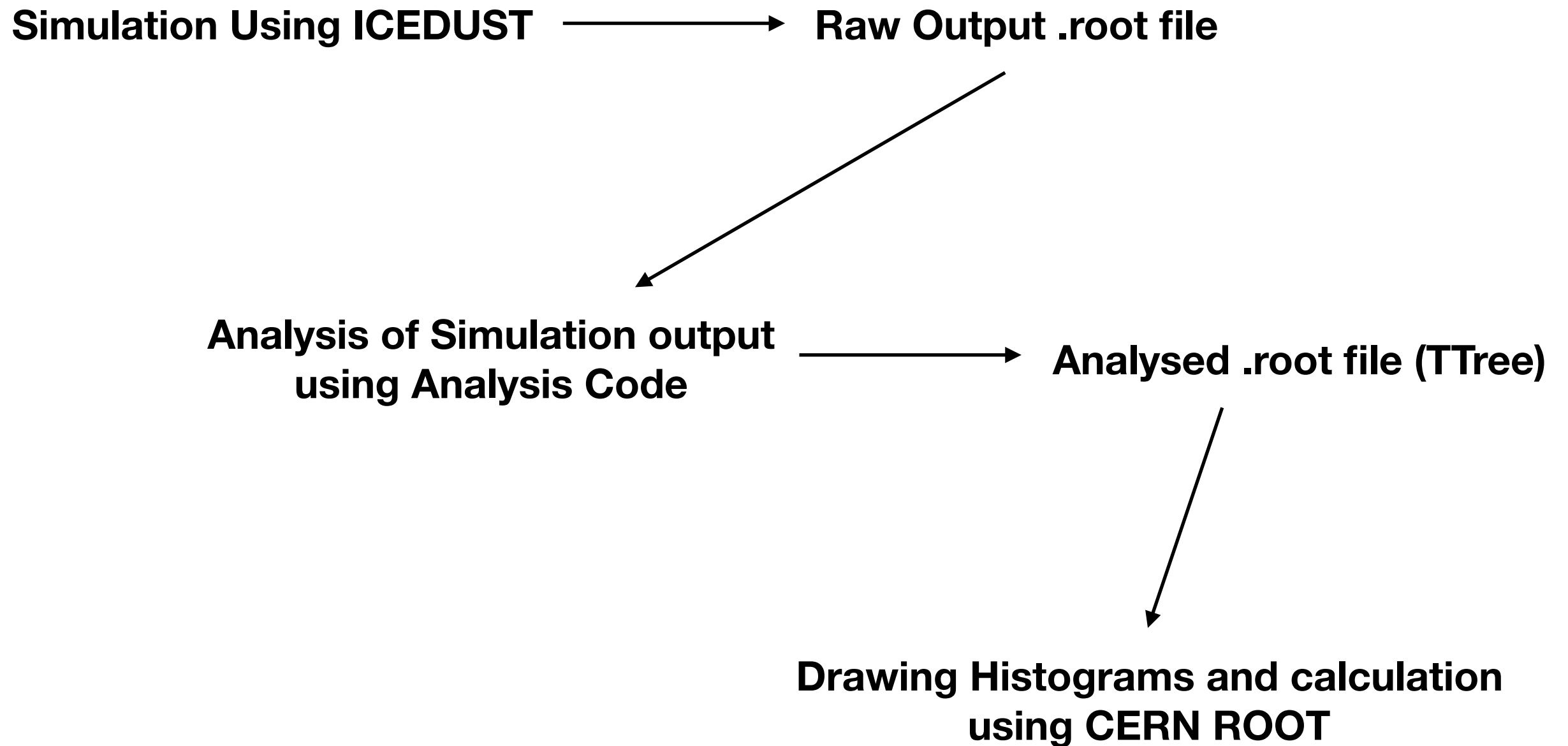
Decay-in-orbit

- DIO Spectrum



From Czernicki's Calculations

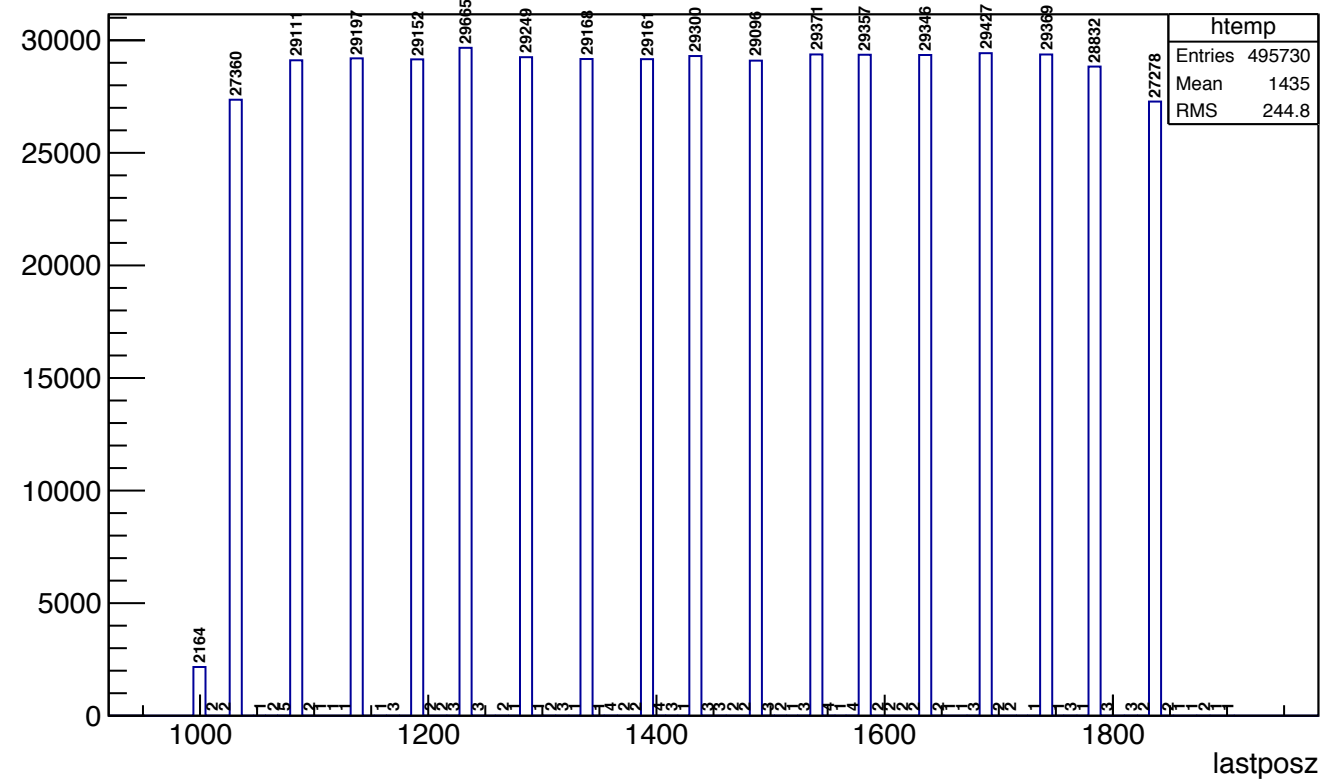
Analysis Workflow



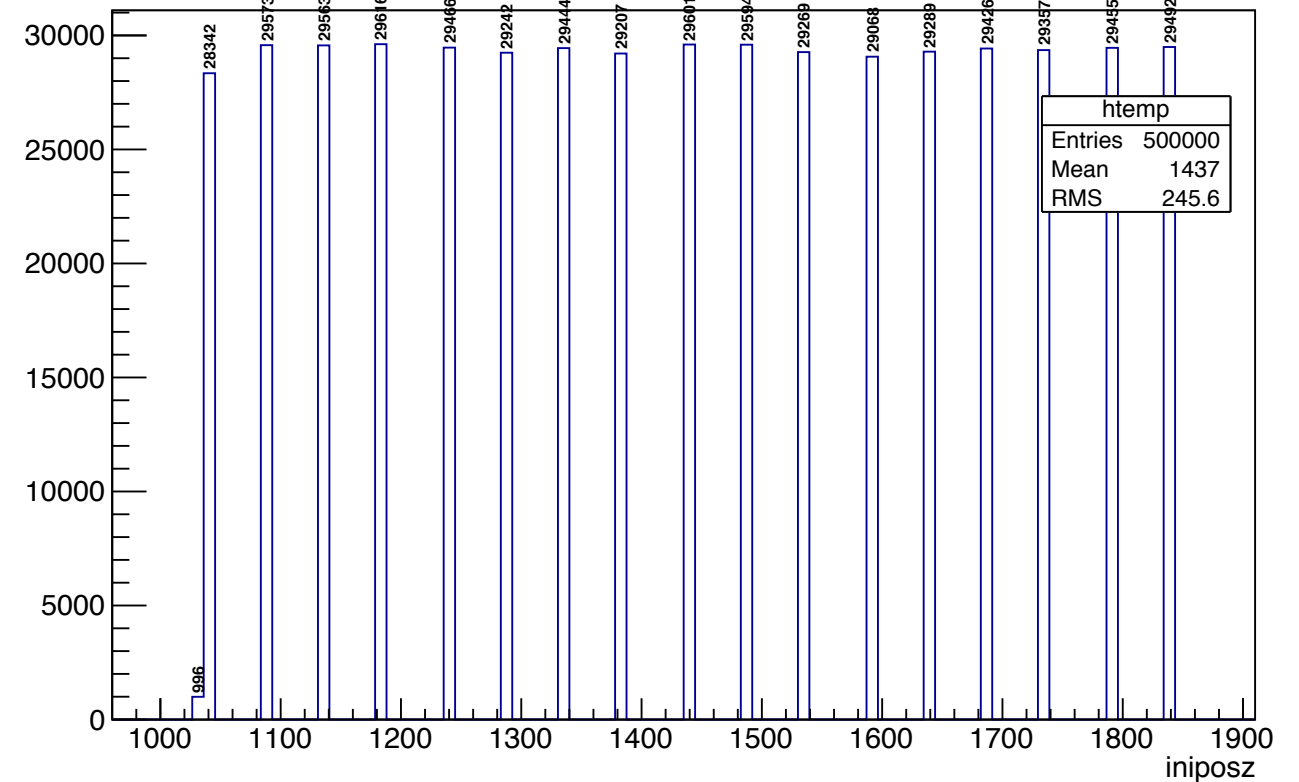
Signal and DIO Simulation

- 500000 Muons uniformly distributed in 17 disks
- Beam Blocker removed
- DIO Blocker removed

The number of stopped muons in each disks

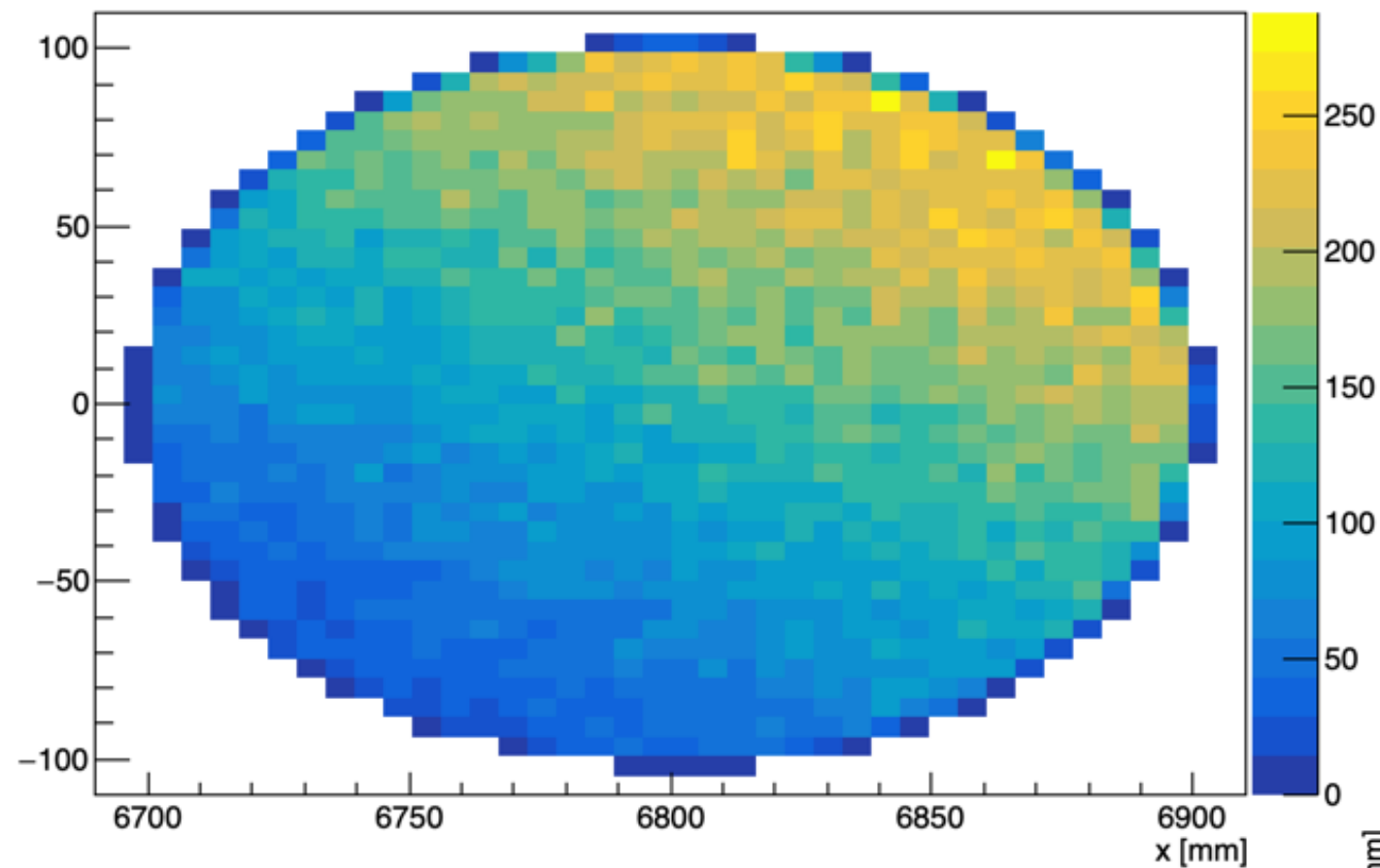


Initial electron with 105MeV

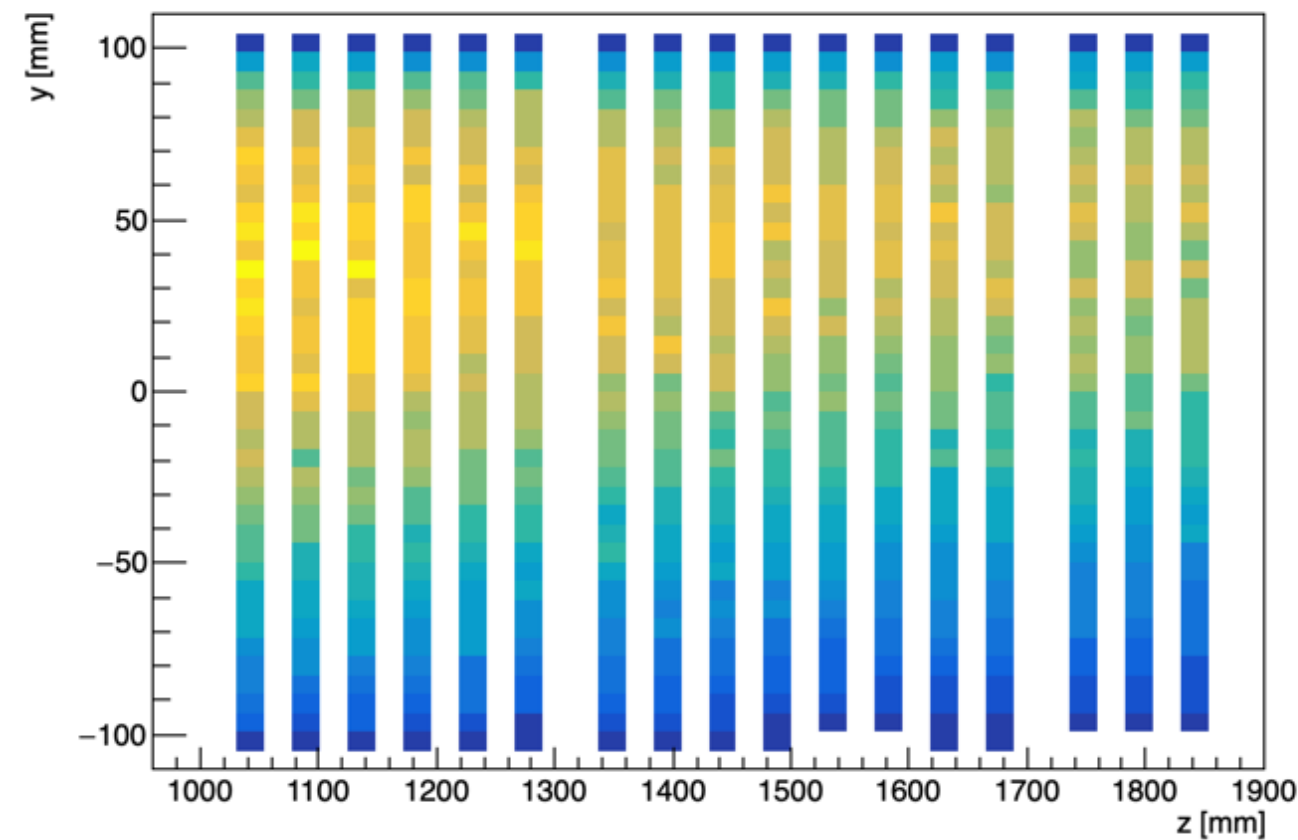


DIO Electrons

Initial Position of DIO Electrons

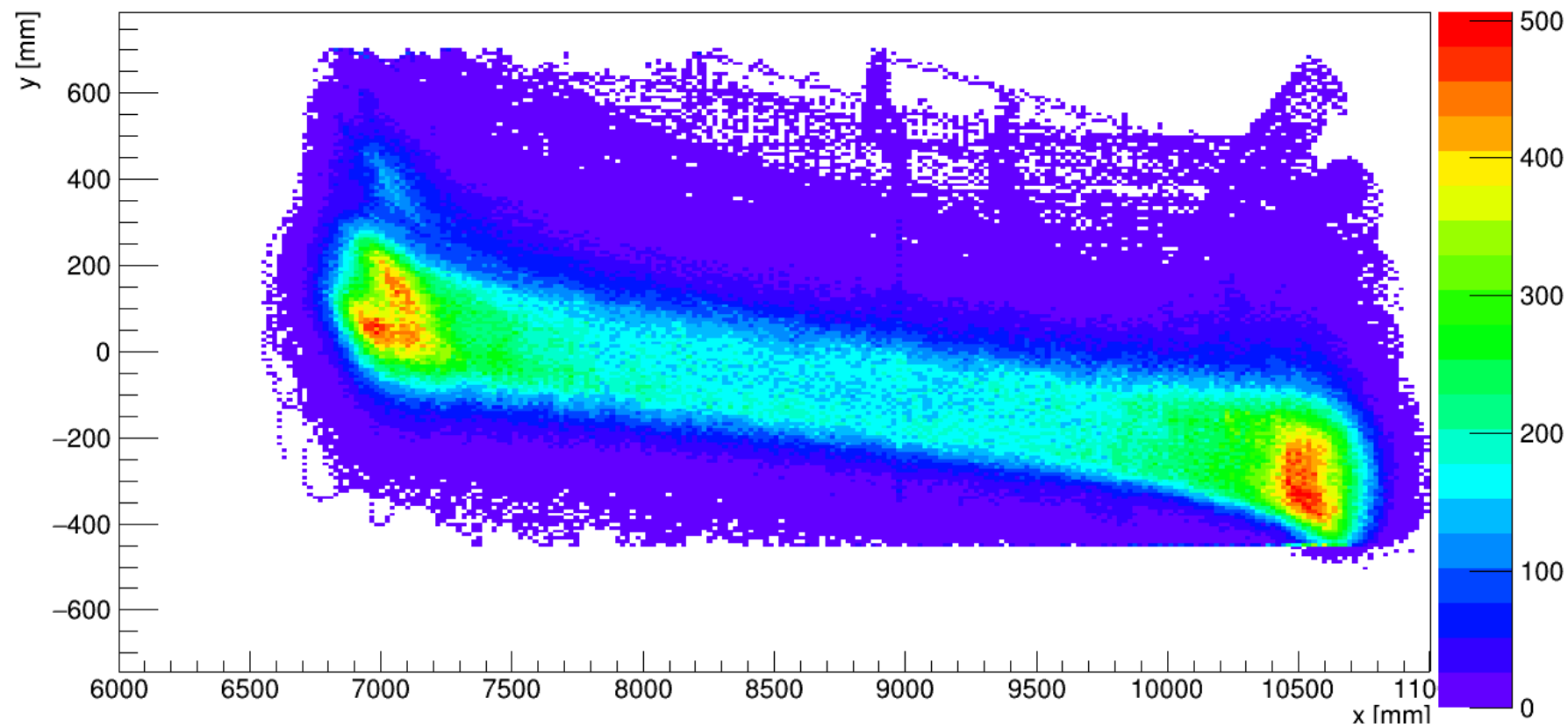


Initial Positio of DIO Electrons

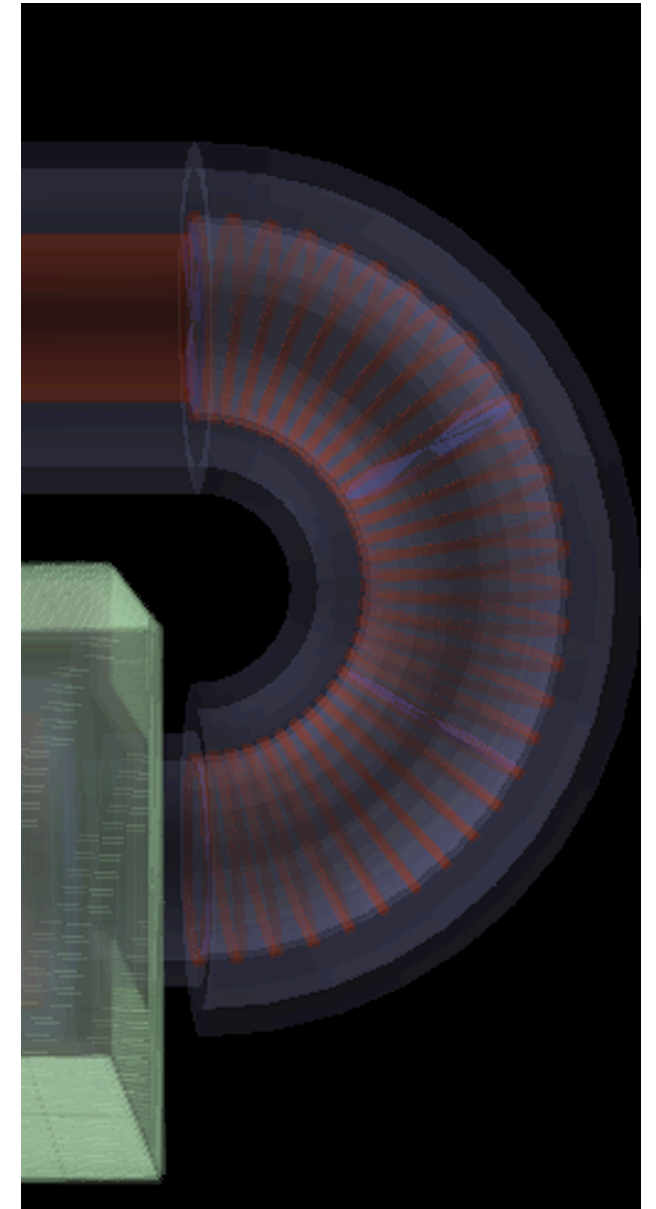
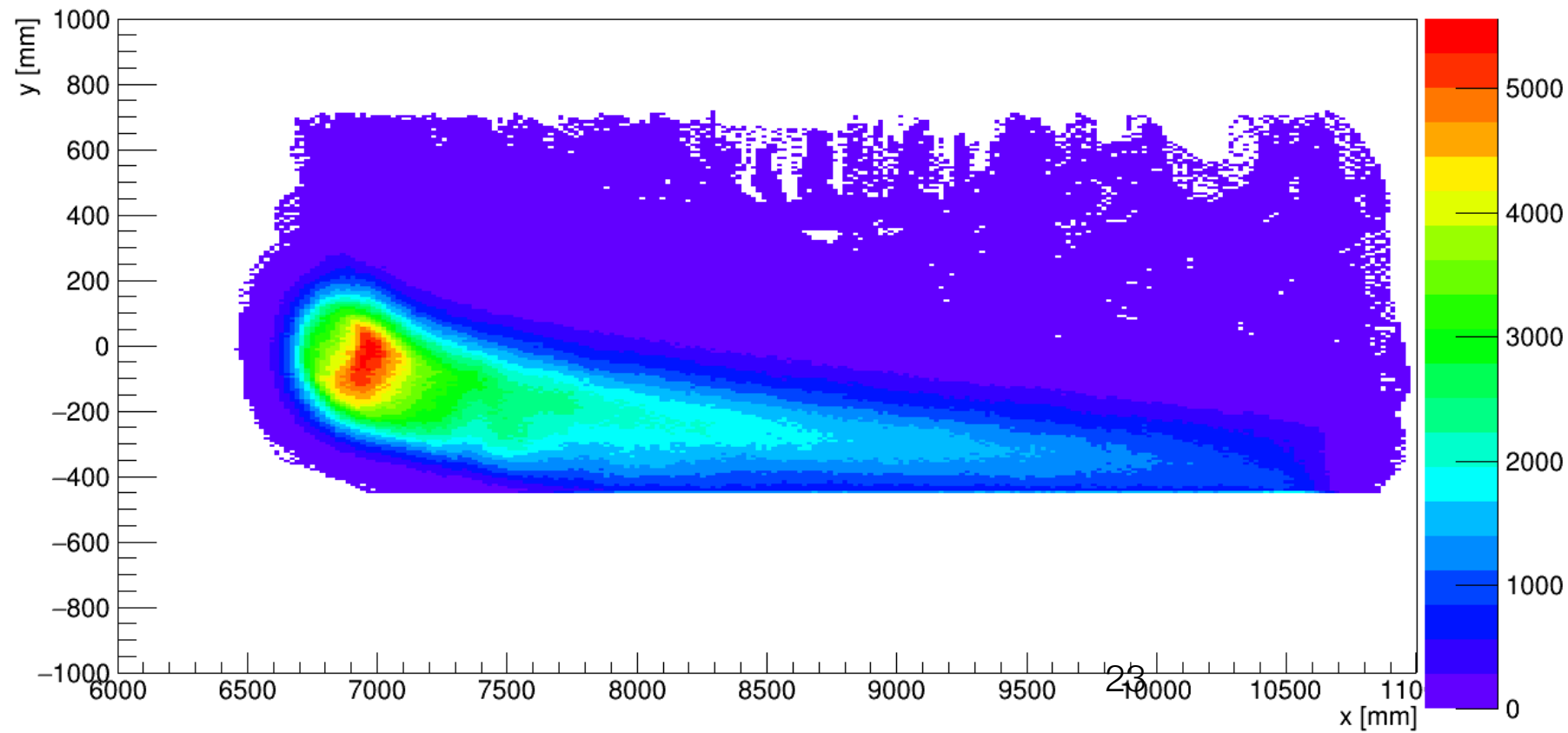


DIO Distribution in Spectrometer

DIO electrons that passed spectrometer

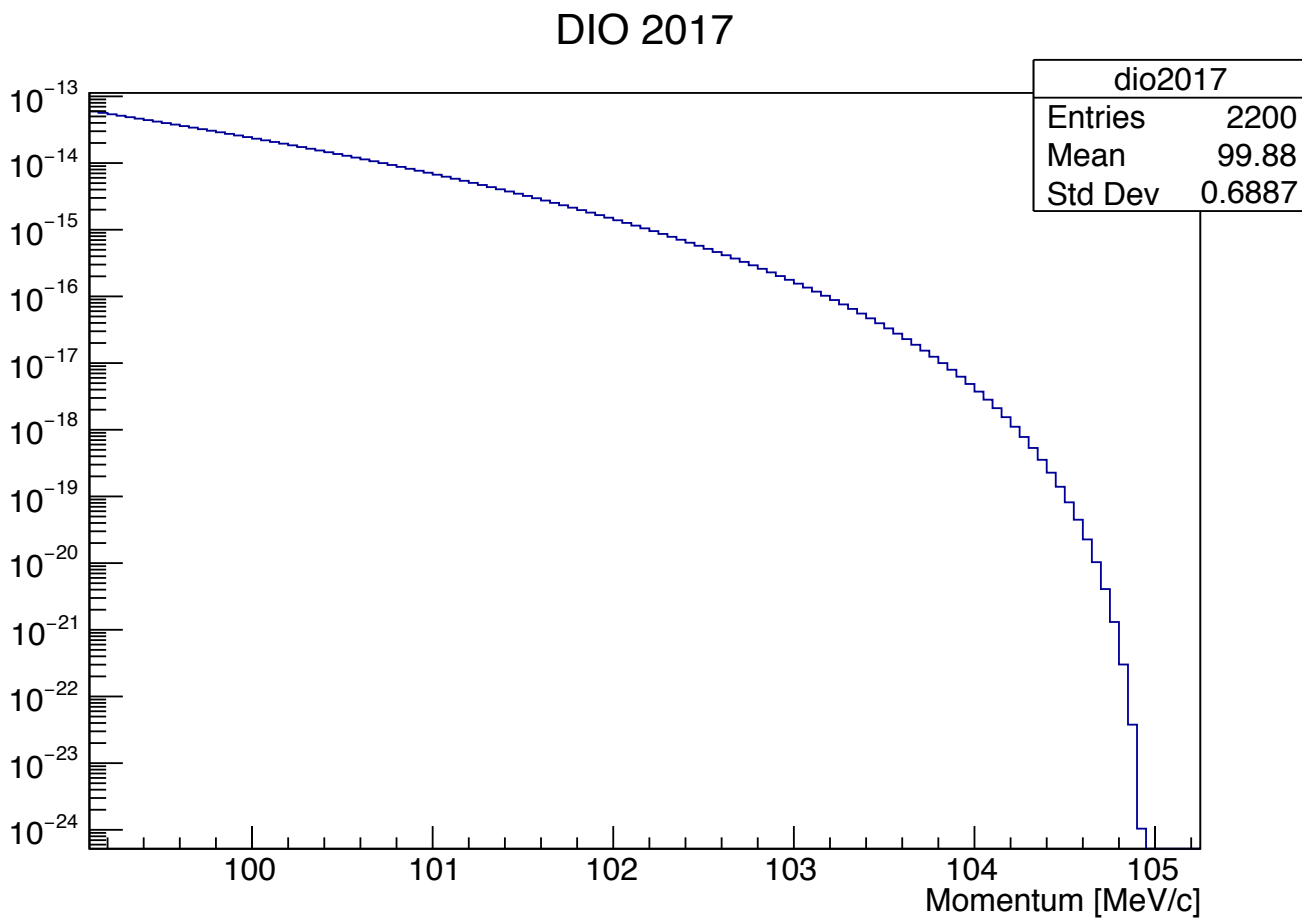


DIO electrons that didn't pass spectrometer

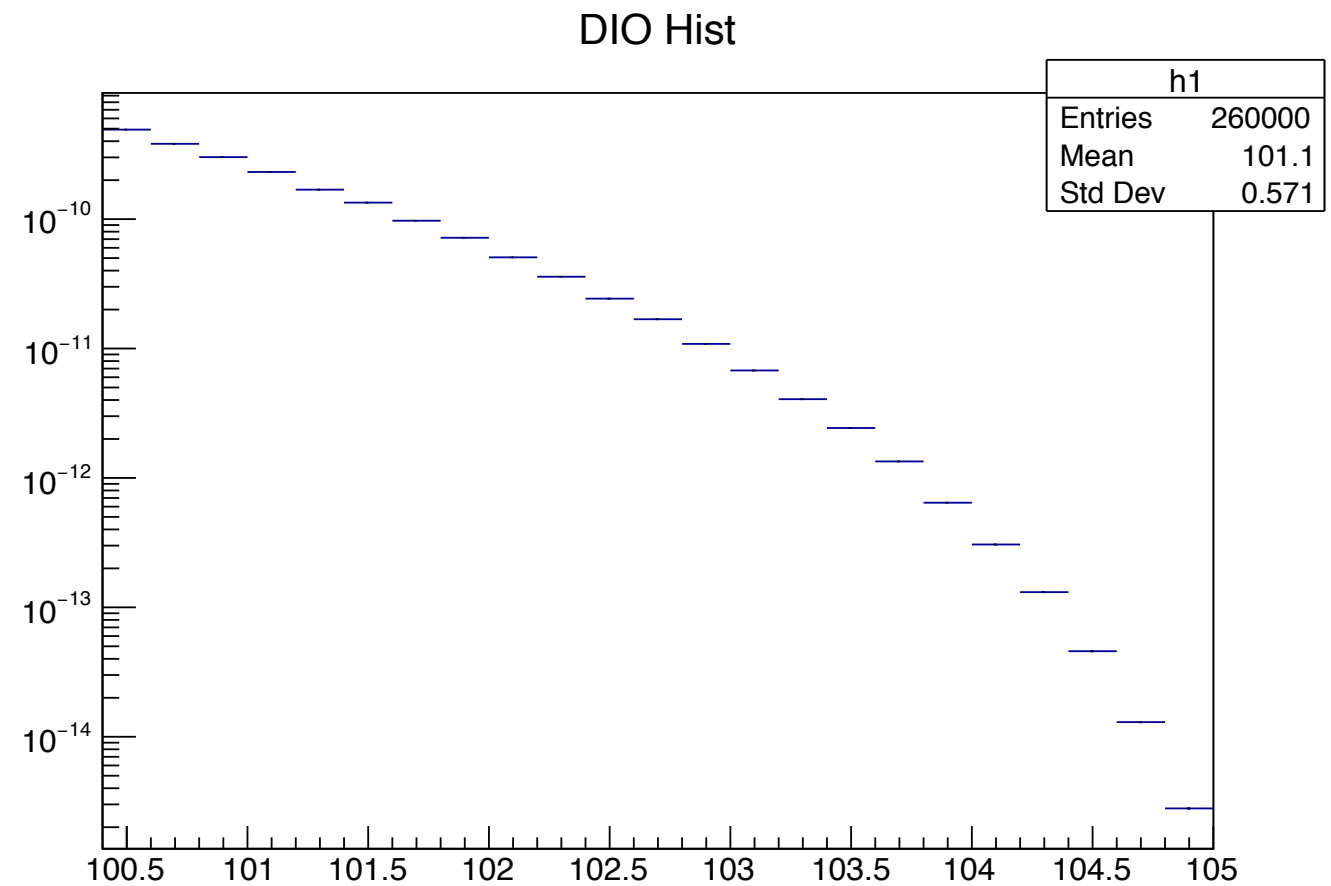


Plots by Wu Chen

DIO Spectrum Analysis



From Czarnecki's Calculations

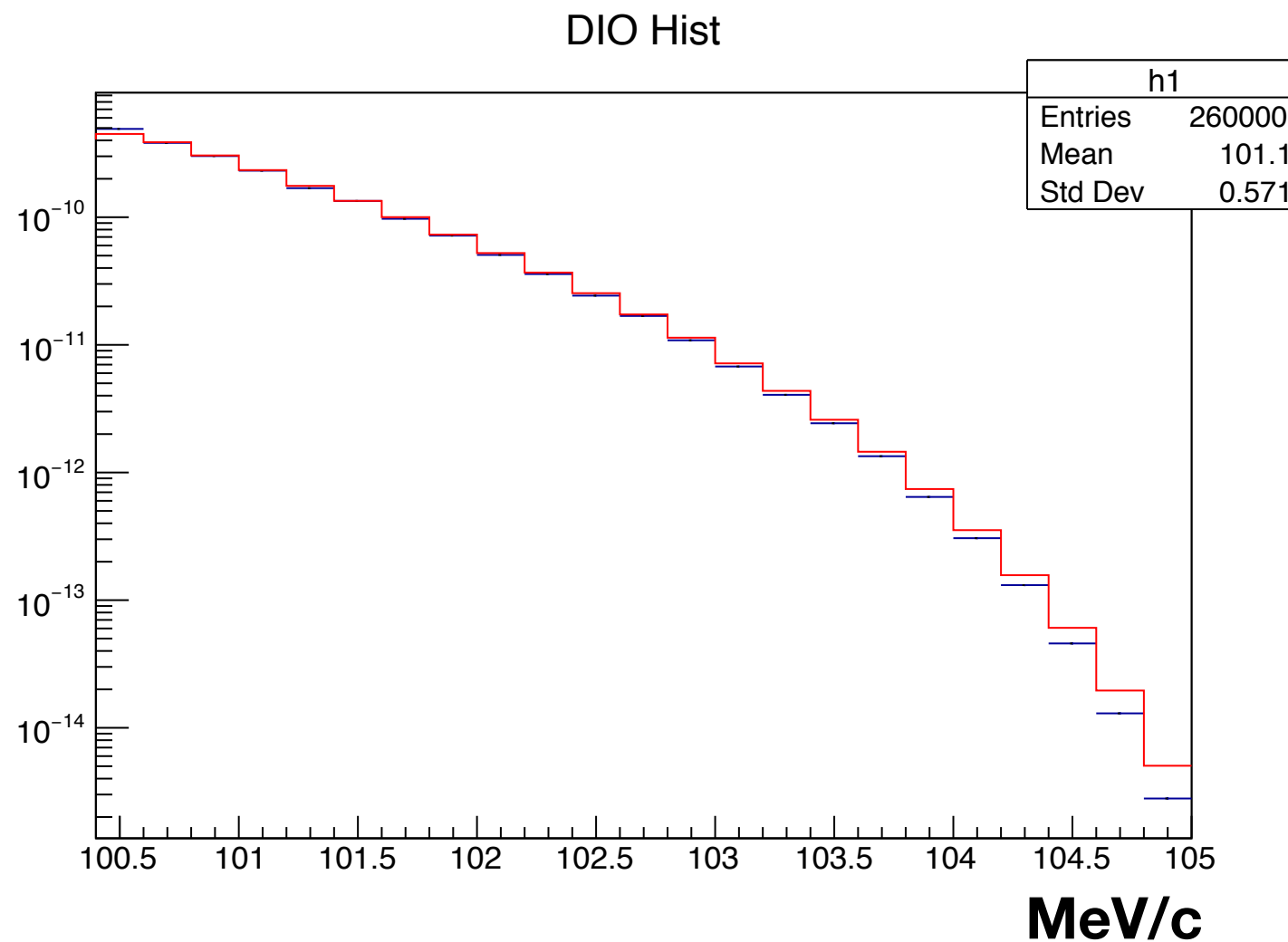


DIO Spectrum recorded in the Straw Tracker

DIO Spectrum with logarithmic scale

Smeared DIO

DIO Spectrum smeared with a Gaussian with sigma of 150 KeV



**Big thanks to Chen Wu
for helping with the code!**

Acceptance and Improvement factors

	this study	Ben's study [4]	CDR (2009) [3]
geometrical acceptance	0.52	0.22	0.22
$A_{\mu \rightarrow e}$ (net signal acceptance)	0.114	0.057	
DIO rejection	6×10^{-4}	4×10^{-5}	

	this study	Ben's study [4]	Improvement factor
N_p	4.3×10^{13}	4.3×10^{13}	
$R_{\mu/p}$	4.8×10^{-3}	1.61×10^{-3}	3
B_{cap}	0.61	0.61	
$A_{\mu \rightarrow e}$	0.114	0.057	2
t_{run}	3×10^7	1.57×10^7	1.9
S.E.S	2.3×10^{-18}	2.6×10^{-17}	11

Sensitivity

	Single event sensitivilty	total POT ($\times 10^{19}$)	Beam time (s)
COMET Phase-II (Ben Krikler 2016)	2.6×10^{-17}	68.3	1.57×10^7
COMET Phase-II CDR2009)	2.6×10^{-17}	85	2×10^7
Mu2e	2.4×10^{-17}	36	6×10^7
COMET Phase-I	3.0×10^{-15}	3.2	1.26×10^7

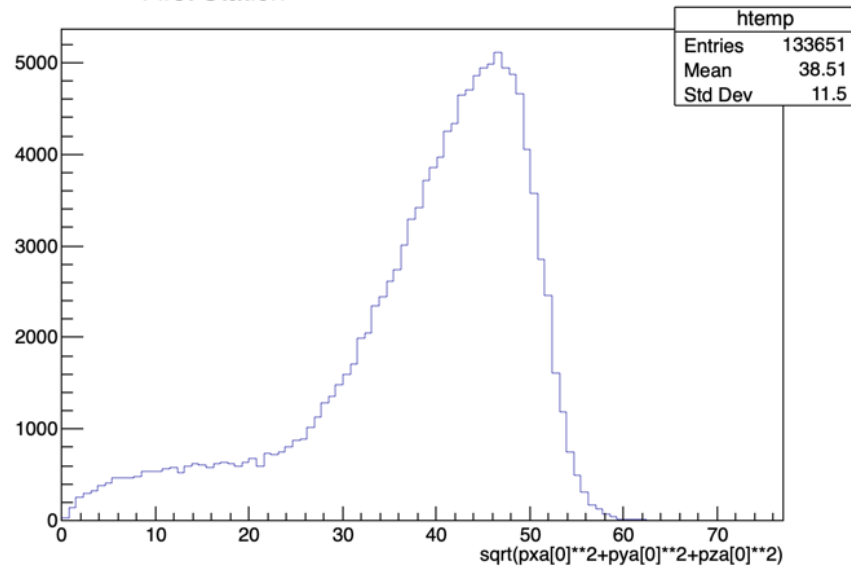
$$\text{SES} = \frac{1}{N_p \cdot R_{\mu/p} \cdot B_{\text{cap}} \cdot A_{\mu \rightarrow e} \cdot t_{\text{run}}}$$

Stopping Rates

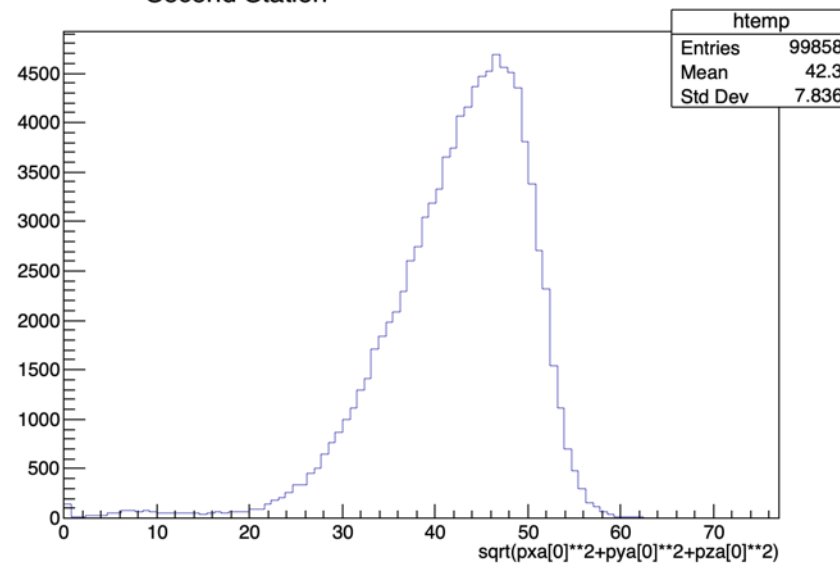
	this study	Ben's study	CDR ((2009)
target radius	4 mm	10 mm	4 mm
target length	400 mm	320 mm	
target location	2 mm downstream	8 mm downstream	
proton beam	1.3 mm(H) $\times$ 1.4 mm(V)	5.8 mm(H) $\times$ 2.9 mm(V)	
muon stop rate	2.63×10^{-3}	1.6×10^{-3}	2.2×10^{-3}

DIO Momentum Distribution on each straw station

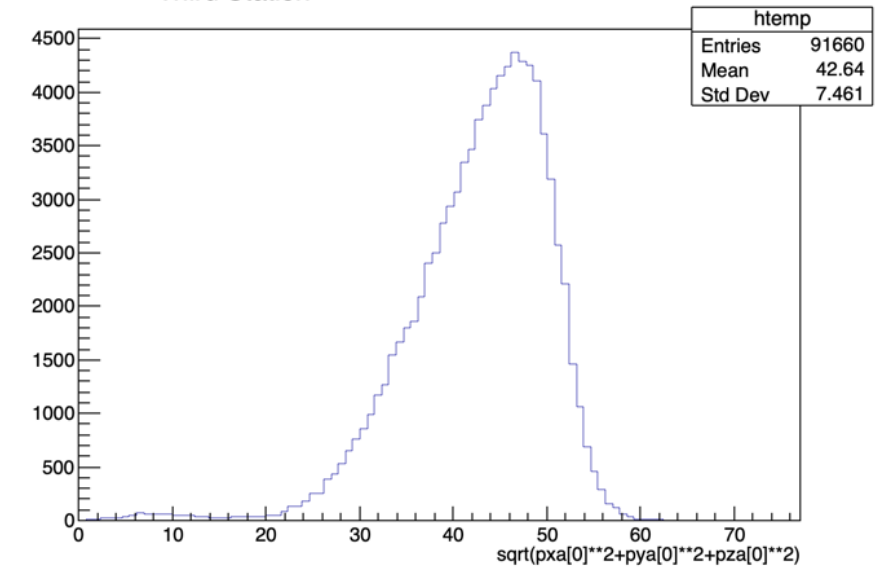
First Station



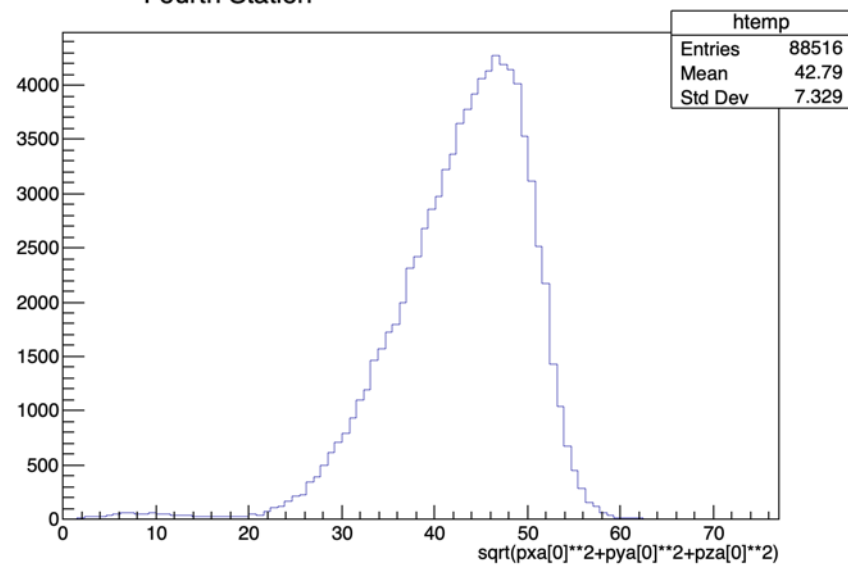
Second Station



Third Station



Fourth Station



Fifth Station

